# <span id="page-0-0"></span>**000 001 002 003** SEE: SEE EVERYTHING EVERY TIME - BROADER LIGHT RANGE IMAGE ENHANCEMENT VIA EVENTS

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# ABSTRACT

Event cameras, with a high dynamic range exceeding  $120dB$ , significantly outperform traditional cameras, robustly recording detailed changing information under various lighting conditions, including both low- and high-light situations. However, recent research on utilizing event data has primarily focused on low-light image enhancement, neglecting image enhancement and brightness adjustment across a broader range of lighting conditions, such as normal or high illumination. Based on this, we propose a novel research question: how to employ events to enhance and adjust the brightness of images captured under broader lighting conditions. To investigate this question, we first collected a new dataset, SEE-600K, consisting of 610,126 images and corresponding events across 202 scenarios, each featuring an average of four lighting conditions with over a 1000-fold variation in illumination. Subsequently, we propose a framework that effectively utilizes events to smoothly adjust image brightness through the use of prompts. Our framework captures color through sensor patterns, uses cross-attention to model events as a brightness dictionary, and adjusts the image's dynamic range to form a broader light-range representation (BLR), which is then decoded at the pixel level based on the brightness prompt. Experimental results demonstrate that our method not only performs well on the low-light enhancement dataset but also shows robust performance on broader light-range image enhancement using the SEE-600K dataset. Additionally, our approach enables pixel-level brightness adjustment, providing flexibility for post-processing and inspiring more imaging applications.

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# 1 INTRODUCTION

**034 035 036 037 038 039 040 041 042 043 044 045 046 047** Every day, from daylight to nighttime, the illuminance varies from about 100,000 *lux* (bright sunlight) to approximately 0.1 *lux* (starlight) [\(Koshel,](#page-11-0) [2012\)](#page-11-0). Maintaining stable imaging under diverse natural lighting conditions is a significant challenge. To achieve this, a series of influential works have emerged, including automatic exposure [\(Bernacki,](#page-10-0) [2020\)](#page-10-0), exposures correction [\(Yuan & Sun,](#page-13-0) [2012\)](#page-13-0), low-light enhancement [\(Li et al.,](#page-11-1) [2021\)](#page-11-1) and high dynamic range (HDR) imaging [\(McCann & Rizzi,](#page-11-2) [2011\)](#page-11-2). However, traditional cameras are limited by their imaging principle of synchronously capturing intensity values across the entire sensor, with a dynamic range of only 60 to 80  $dB$  [Hasinoff et al.](#page-10-1) [\(2016\)](#page-10-1); [Rebecq et al.](#page-12-0) [\(2019\)](#page-12-0). Consequently, these traditional methods find it difficult to capture imaging information under a wide range of lighting conditions at the input [\(Gehrig & Scaramuzza,](#page-10-2) [2024;](#page-10-2) [Gallego](#page-10-3) [et al.,](#page-10-3) [2020\)](#page-10-3). If the exposure is inaccurate - over and under exposures - traditional cameras lose the potential to restore images under complex lighting conditions due to limited bits-width and noise. Unlike traditional cameras, event cameras [Gallego et al.](#page-10-3) [\(2020\)](#page-10-3) asynchronously record pixel-level changes in illumination, outputting the direction of intensity change (positive or negative) at each pixel with extremely high dynamic range  $(120 \text{ dB})$ , which far exceeds the capability of traditional cameras in capturing various lighting intensity.

**048 049 050 051 052 053** Research leveraging the events for image brightness enhancement can be divided into three categories. (1) event-based image reconstruction, which aim to reconstruct images only from events. However, these methods [\(Rebecq et al.,](#page-12-0) [2019;](#page-12-0) [Stoffregen et al.,](#page-12-1) [2020;](#page-12-1) [Wang et al.,](#page-12-2) [2024\)](#page-12-2) rely solely on events, facing uncertainties during reconstruction, and the events usually contain heavy noise, which leads to color distortion and limited capabilities of generalization. (2) event-guided HDR imaging [\(Cui et al.,](#page-10-4) [2024;](#page-10-4) [Yang et al.,](#page-13-1) [2023;](#page-13-1) [Messikommer et al.,](#page-11-3) [2022\)](#page-11-3) , which aims to employ events to extend the dynamic range of images or video to match human vision. However, synthesizing HDR images as ground truth

<span id="page-1-1"></span>**054** SDE: Brightness Distribution  $I_i$ **055** 乳腺 ICL  $0.5$ Low Light  $I<sub>o</sub>$ Previous **056**  $\cal E$ Normal Light Methods (c) (a) **057** Inputs (Image  $I_i$  & Event  $E$ ) EvLowLight Output  $I_o$ **058**  $0.2$   $0.3$ <br>Brightness (0 to 1)  $0.1$  $0.4$  $0.5$ *Only Low-Light Exposure No control over brightness* EvLight, *et.al.* **059 060** SEE: Brightness Distribution -133 牙形  $I_i$ Brightness 0.5 **061 ICLR**  $I<sub>o</sub>$ SEE Net 15 **062**  $\cal E$ (*Ours*) Low Light<br>Normal Light nt in **063** 10 (b) (d) 27. IP High Light  $B$  Brightness is **064** controlled by  $\bar{B}$ **065**  $f_b(I_o) \rightarrow B$ 0 **1** 1 Inputs (Image  $I_i$  & Event  $E$ ) **066**  $0.2$  $0.4$  $0.6$  $0.8$  $1.0$  $0.0$ *Various of Exposure* Brightness Prompt Brightness (0 to 1) Outputs  $I<sub>a</sub>$ **067**

<span id="page-1-0"></span>Figure 1: (a) and (b): Brightness distributions of the SDE dataset (0∼0.45, low to normal light) and our SEE-600K dataset (0∼1, a broader light range). (c): Previous methods [\(Liang et al.,](#page-11-4) [2024;](#page-11-4) [2023\)](#page-11-5) directly map low-light images to normal-light images. (d): Our SEENet accepts inputs across a broader brightness range and adjusts output brightness through prompts.  $f<sub>b</sub>$  refers to the function that calculates the brightness of an image.

**073 074 075 076 077 078 079 080 081 082 083 084 085 086 087 088 089 090 091 092** is difficult. [\(Cui et al.,](#page-10-4) [2024\)](#page-10-4) introduced the first real-world dataset containing paired color events, low dynamic range, and HDR images, with only includes 1,000 HDR images. [\(Messikommer et al.,](#page-11-3) [2022\)](#page-11-3) used nine images with different exposures to synthesize an HDR image as the ground truth and utilized multi-exposure frames and events as inputs to generate an HDR image. While HDR imaging aims to expand dynamic range, collecting HDR datasets is difficult, and these methods have not been evaluated for tasks like low-light enhancement or high-light restoration [\(Tursun et al.,](#page-12-3) [2015;](#page-12-3) [Jayasuriya](#page-10-5) [et al.,](#page-10-5) [2023\)](#page-10-5). (3) event-guided low-light enhancement [\(Liang et al.,](#page-11-4) [2024;](#page-11-4) [2023;](#page-11-5) [Liu et al.,](#page-11-6) [2023;](#page-11-6) [Jiang](#page-10-6) [et al.,](#page-10-6) [2023\)](#page-10-6), which is designed to adjust low-light images to normal-light conditions through brightness adjustment and noise reduction. [Liang et al.](#page-11-4) [\(2024\)](#page-11-4) represents the latest research and proposed the first event-based low-light image enhancement dataset, SDE (see Fig. [1](#page-1-0) (a)). Prior to this, [Liang et al.](#page-11-5) [\(2023\)](#page-11-5); [Liu et al.](#page-11-6) [\(2023\)](#page-11-6); [Jiang et al.](#page-10-6) [\(2023\)](#page-10-6) explored using motion information from events and employed varying neural networks to improve the mapping from low-light images to normal-light ones, as shown in Fig.  $1$  (c). However, these strategies only focus on the improvement of mapping ability for low-light inputs, limiting their capacity to adjust brightness across a broader range of lighting conditions, *e.g.*, normal or high-light images. Furthermore, due to the uncertainty in the standard for normal-light image collection—as the normal-light images are relative to low-light images (as shown in Fig[.1](#page-1-0) (a))—these methods introduce ambiguity during the training process because they can only map low-light images to normal-light ones based on a single set of low- and normal-light data pairs captured per scene. Overall, current research focuses on low-light enhancement, neglecting image enhancement and processing under a wider range of lighting conditions. Therefore, *how to use events to enhance and adjust the brightness of images across a broader range of lighting conditions* becomes a more worthwhile research question.

**093 094 095 096 097 098** To address this novel research question, we first formulate the imaging model for brightness adjustment (Sec[.3\)](#page-2-0) and define the learning task. We aim to perceive lighting information from events, utilizing brightness prompts to convert this lighting information into images with a specific brightness. In doing so, other image quality aspects (like sensor patterns, noise, color bias, and so on) are taken into consideration.

**099 100 101 102 103 104 105 106 107** To realize our proposed task, we first collecte a new dataset by emulating each scene in different lighting conditions, covering a broader luminance range (Sec[.4\)](#page-3-0), as shown in Fig[.1](#page-1-0) (b) and (d). By capturing multiple lighting conditions per scene, we enable mappings across diverse illumination scenarios, providing rich data for model training. To tackle the challenges of spatio-temporal alignment of video and event streams under various lighting conditions, we design a temporal alignment strategy relying on programmable robotic arms and inertial measurement unit (IMU) sensors. As a result, we obtain a temporal registration error up to one millisecond and a spatial error at the sub-pixel level (∼ 0.3 *pixel*). Finally, we build a large-scale and well-aligned dataset containing 202 scenes, each with 4 different lighting conditions, summing up to 610,126 images and the corresponding event data. We term this dataset as SEE-600K, which supports learning the mappings among multiple lighting conditions.

<span id="page-2-1"></span>**108 109 110 111 112 113** Building on the SEE-600K dataset, we propose a compact and efficient framework, SEE-Net, for the proposed new tasks (Sec. [5\)](#page-5-0). An event-aware cross-attention is used to enhance image brightness, and the brightness-related prompt is introduced for controlling the overall brightness. This approach effectively captures and adjusts lighting across a broader range of illumination conditions, providing flexibility and precise control during inference. Despite of the advantage of performance, SEE-Net still remains effective, compact, and lightweight with only 1.9 *M* parameters.

**114 115 116 117 118** Our method has been evaluated on two real-world datasets, SDE [\(Liang et al.,](#page-11-4) [2024\)](#page-11-4) and SEE-600K. Quantitative results demonstrate that our framework fits well to a broader range of lighting conditions (Sec. [6\)](#page-6-0). Furthermore, our framework allows for smooth brightness adjustment, providing precise exposure control. Therefore, this flexibility significantly improves post-processing capabilities and enables potential applications in advanced imaging and processing tasks.

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# 2 RELATED WORKS

**122 123 124 125 126 127 128 129 130 131 132 133 134 135 136 137 138 139 140 141 142 143 144 145 146** Frame-based: These brightness enhancement methods aim to improve image quality under chal-lenging illumination conditions. Retinexformer [\(Cai et al.,](#page-10-7) [2023\)](#page-10-7) and other Retinex-based frameworks [\(Zhang et al.,](#page-13-2) [2021;](#page-13-2) [Wu et al.,](#page-13-3) [2022;](#page-13-3) [Fu et al.,](#page-10-8) [2023\)](#page-10-8) decompose reflectance and illumination with complex training pipelines. Other approaches, *e.g.*, structure-aware models [\(Xu et al.,](#page-13-4) [2023b;](#page-13-4) [Wang et al.,](#page-12-4) [2023c\)](#page-12-4), utilize edge detection or semantic-aware guidance to achieve sharper and more realistic results. Exposure correction strategies [\(Afifi et al.,](#page-10-9) [2021;](#page-10-9) [Panetta et al.,](#page-12-5) [2022;](#page-12-5) [Ma et al.,](#page-11-7) [2020\)](#page-11-7) target both overexposed and underexposed areas, leveraging multi-scale networks or perceptual image enhancement frameworks to synthesize correctly exposed images. However, the reliance on RGB frames with limited bit depth, limits the adaptability to dynamic lighting conditions, making it difficult to handle a broader range of lighting scenarios. **Event-based:** These methods focus on reconstructing images or videos exclusively from event data. For instance, [Duwek et al.](#page-10-10) [\(2021\)](#page-10-10) introduced a two-phase neural network combining CNNs and SNNs, while [Pan et al.](#page-12-6) [\(2019\)](#page-12-6) proposed the event-based double integral model to generate videos. [Stoffregen et al.](#page-12-1) [\(2020\)](#page-12-1) enhanced event-based video reconstruction by introducing the new dataset. Additionally, [Liu & Dragotti](#page-11-8) [\(2023\)](#page-11-8); [Wang et al.](#page-12-2) [\(2024\)](#page-12-2) developed a model-based deep network to improve reconstructed video quality. However, these event-based approaches face challenges due to event data noise, often leading to color distortion and limited generalization. **Event-guided:** These works are centered on enhancing images captured in low-light conditions. E.g., [Zhang et al.](#page-13-5) [\(2020\)](#page-13-5) and [Liu et al.](#page-11-9) [\(2024\)](#page-11-9) recovered lost details in low-light environments by reconstructing grayscale images. Similarly, [Liang et al.](#page-11-5) [\(2023\)](#page-11-5) and [Liu et al.](#page-11-6) [\(2023\)](#page-11-6) improved low-light video enhancement by leveraging motion information from events to enhance multi-frame videos and integrating spatiotemporal coherence. Furthermore, [Jin et al.](#page-11-10) [\(2023\)](#page-11-10) and [Jiang](#page-10-6) [et al.](#page-10-6) [\(2023\)](#page-10-6) utilized events to recover structural details and reconstruct clear images under near-dark situations. Most notably, [Liang et al.](#page-11-4) [\(2024\)](#page-11-4) introduced the first large-scale event-guided low-light enhancement dataset, which is significant for the development of this field. While these methods use events for brightness changes and structural recovery in low-light conditions, they are limited to enhance low-light images with single mapping and cannot handle brightness adjustments across a broader range of lighting conditions, including normal- and high-light.

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# <span id="page-2-0"></span>3 PRELIMINARIES AND NEW TASK DEFINITION

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**151 152 153 154 155 156** In this section, we formalize the physical model underlying our approach to enhance and adjust image brightness across a broader range of lighting conditions using events. Imaging is fundamentally the process of capturing the radiance of a scene, represented as a radiance field  $L(t)$  varying over a preset slot t. The illuminates of light in daily life span a vast range, from 0.1 *lux* (starlight) to 1e6 *lux* (direct sunlight). The goal of brightness adjustment is to recover or estimate  $L(t)$  and tone-map it into an image that is visually suitable for human perception.

**157 158 159 160 161** Traditional cameras record light signals through exposure [\(Mendis et al.,](#page-11-11) [1997\)](#page-11-11). This voltage is influenced by the Gaussian noise  $N = \mathcal{N}(\mu, \sigma^2)$  ( $\mu$  is the mean and  $\sigma^2$  is the variance), and the photon shot noise  $P = \mathcal{P}(k)$ , where  $k \propto L(t)$  is the number of photons, proportional to light intensity. In low-light conditions, Gaussian noise dominates, while in high-light conditions, photon shot noise becomes more significant. These noises influence the final value in the RAW image, simply represented as  $I_{\text{raw}} \approx \mathcal{Q}(L(t) + P + N)$ , where  $\mathcal Q$  is the quantization function that converts the

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**175 176 177** Figure 2: (a) data collection setup: Universal Robots UR5e arm replicates precise trajectories with an error margin of  $0.03mm$ . (b) **IMU data registration:** *b* (1) shows unregistered IMU data, while *b* (2) displays registered data after timestamp alignment. (c) EVS outputs with different filters: *f1* to *f4* demonstrate the different ND filters, depicting various lighting levels.

**179 180 181 182 183 184 185 186 187 188 189** continuous voltage into discrete digital signals, typically ranging from 8 to 12 bits. The shape of the image  $I_{\text{raw}}$  is  $H \times W \times 1$ , where H and W are the image resolution. The RAW image is then further processed through image signal processing (ISP) fisp, which includes multiple steps *e.g.*, denoising, linear and non-linear transformations, resulting in a RGB image as  $I_{\text{rgb}} = f_{\text{isp}}(I_{\text{raw}})$ , with the shape of  $H \times W \times 3$ . An accurate image exposure procedure recovers  $I_{rgb}$  corresponding to  $L(t)$ , up to a high degree meeting the following three characteristics: (1) accurate exposure: The mean value of  $I_{\text{rgb}}$  falls within the range [0.4, 0.7] [\(Mertens et al.,](#page-11-12) [2009\)](#page-11-12). (2) noise-free: The influence of N and P is suppressed to a visual-acceptable level. (3) color neutrality: The gray levels calculated from the RGB channels should be consistent [\(Buchsbaum,](#page-10-11) [1980\)](#page-10-11). However, traditional cameras sometimes fail to capture sufficient details in extreme-lighting scenes. Under such low-light conditions, images may lack visible details and be contaminated by noise, while in high-light conditions, images may suffer from oversaturation, losing texture and edge information.

**190 191 192 193 194 195 196 197 198 199 200** Event cameras asynchronously detect illumination changes at each pixel, making them ideal for capturing scenes with extreme or rapidly changing lighting conditions [\(Gallego et al.,](#page-10-3) [2020\)](#page-10-3). The event stream's outputs are formatted as 4 components:  $(x, y)$  (pixel coordinates), t (timestamp), and  $p \in \{+1, -1\}$  (polarity, indicating light intensity increase or decrease). Events are triggered when the change in illumination exceeds a threshold  $C (\Delta L = \log(L(t)) - \log(L(t - \Delta t))$  where  $|\Delta L| > C$ ). We jointly leverage the complementary information from an image  $I_{rgb}$  and its corresponding events E to recover a high-quality well-illuminated image  $I_{rgb}$  that accurately represents the scene radiance  $L(t)$ , while also allowing for adjustable brightness. To achieve this, we introduce a brightness prompt B that controls the overall brightness of the output image. This allows us to map the  $L(t)$  into an image that is optimally exposed for human observation. Our task setting can thus be formulated as Eq. [1,](#page-3-1) where  $f_{see}$  is our proposed model, as shown in Fig. [1.](#page-1-0)

<span id="page-3-1"></span>
$$
f_{see}(I_{rgb}, E, B) \to \hat{I_{rgb}}.
$$
 (1)

**202 203 204 205 206 207 208** This formulation has two advantages: (1) robust training: By inserting the brightness prompt  $B$ during training, we can decouple the model from biases in the training data with specific brightness level, enabling the model to generalize better over illuminates domain. (2) flexible inference: During inference, the prompt B can be set to a default value ( $e.g., B = 0.5$ ) to produce images with general brightness, or be adjusted to achieve different brightness levels, providing flexibility for applications requiring specific exposure adjustments or artistic effects. *Due to space limitations, please refer to the supplementary material for more details of this section.*

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# <span id="page-3-0"></span>4 DATASET COLLECTION

**212 213 214 215** In this section, we introduce the SEE-600K dataset, designed to contain *(1) multiple lighting conditions*, *(2) complex motion trajectories* and *(3) spatio-temporal alignment*. Unlike the state-of-the-art SDE dataset [\(Liang et al.,](#page-11-4) [2024\)](#page-11-4), we capture data across multiple lighting conditions. Most importantly, SEE-600K is nearly 20 times larger than the SDE dataset, providing a stronger foundation for training models with better generalization.

<span id="page-4-1"></span>

<span id="page-4-0"></span>Figure 3: (a) registration process: Illustration of the multi-level registration process, showing how trajectories, S and T, at various levels are iteratively aligned. (b) two trajectories: Example of two aligned images captured along two trajectories. (c) pixel distance change: Temporal distance of pixel between two registered videos, showing a mean alignment error of 0.2957 pixels over time.

**234 235 236 237 238 239 240 241 242 243 244 245 246 247 248** (1) multiple lighting conditions: Our approach is based on the principle that lighting transitions continuously from low to high intensity. Unlike previous datasets [\(Liang et al.,](#page-11-4) [2024;](#page-11-4) [Wang et al.,](#page-12-7) [2021\)](#page-12-7), which captured only a *single* pair of low-light and normal-light conditions, we focus on *multiple* samples. To cover a broader lighting range, we record an average of four videos per scene, using neutral density (ND) filters at three levels  $(1/8, 1/64, 1/1000)$  and one without a filter. We also adjust the aperture and exposure settings to capture each scene under diverse lighting conditions. (2) complex motion trajectories: We employ the Universal Robots UR5e robotic arm, which can provid high stability and repeat the same non-linear trajectory with an error margin of **0.03 mm** [\(Liang et al.,](#page-11-4) [2024;](#page-11-4) [Brey et al.,](#page-10-12) [2024\)](#page-10-12), allowing us to capture multiple videos with spatial consistency, as exhibited in Fig.  $3$  (a). (3) spatio-temporal alignment: While the robotic arm guaranteed spatial alignment, asynchronous control over the camera's start and stop times inevitably introduced timing deviations. To resolve this, we propose an IMU-based temporal alignment algorithm, as shown in Fig. [3](#page-4-0) (b). IMU streams synchronized to events and video with microsecond timestamps in the DVS346 camera. Additionally, the IMU stream depends only on motion trajectory and enjoys a temporal resolution of 1000 Hz. Based on this, our algorithm achieves precise temporal alignment, ensuring synchronization across the entire dataset, as displayed in Fig.  $3$  (c).

**249 250 251 252 253 254 255 256 257 258 259 260 261 262** Temporal IMU Registration Algorithm: We propose an IMU data registration algorithm that aligns the source sequence S and target sequence T by finding the optimal bias b and matching length  $l$ to minimize the  $L_1$  distance between them. Given the high resolution of IMU data at  $1000Hz$ , an exhaustive search for the optimal bias is computationally infeasible. To address this, we introduce a multi-level iterative strategy. First, we denoise the IMU data using a Kalman filter [\(Mirzaei &](#page-12-8) [Roumeliotis,](#page-12-8) [2008\)](#page-12-8). Then, the average pooling is utilized to reduce the sequences to two additional levels, Level-1 ( $S_1$ ,  $T_1$ ) and Level-2 ( $S_2$ ,  $T_2$ ), as shown in Fig. [3](#page-4-0) (a)- $\Phi$ 2. This reduces computational complexity while preserving essential alignment features. The window size is chosen based on our video durations, which ranges from 10 to 120 seconds. We perform a coarse search for the optimal bias b and matching length l at the lowest resolution (Level-2). The results from this level serve as center points for finer searches at higher resolutions. Specifically, the bias and length identified at each level guide local searches at the next level up, as displayed in Fig. [3](#page-4-0) (a)- $\Im(\mathfrak{A})\Im(\mathfrak{A})$ . At Level-1 and the original data level (Level-0), we only need to search locally around these center points. This hierarchical approach efficiently achieves high matching accuracy with significantly reduced computational effort.

**263 264 265 266 267 268 269** Spatial-Temporal Alignment Evaluation: To evaluate the accuracy of our IMU registration algorithm, we capture the same scene twice under identical lighting conditions, as illustrated in Fig. [3](#page-4-0) (b). We assess the alignment metric between the two image sequences by calculating the pixel-level distance at the corresponding timestamp. *Alignment Metric*: For each image pair, we extract keypoints using SIFT [\(Lowe,](#page-11-13) [2004\)](#page-11-13) and then employ the FLANN matcher [\(Muja & Lowe,](#page-12-9) [2009\)](#page-12-9) to find matching keypoints between the two images. Based on these matched keypoints, we compute the affine transformation matrix using RANSAC [\(Fischler & Bolles,](#page-10-13) [1981\)](#page-10-13). This transformation is subsequently applied to each pixel, allowing us to calculate the displacement distance for every pixel. Finally, the

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**287 288 289 290** Figure 4: Overview of our proposed framework, called SEE-Net, which is composed of four stages: (a) Inputs and Preprocessing, (b) Encoding, (c) Decoding, and (d) Loss Function. This framework takes as input an image captured under a wide range of lighting conditions, along with its corresponding events. The output is a brightness-controllable image, where the brightness is guided by the brightness prompt  $B$ , enabling flexible pixel-level adjustment during inference.

<span id="page-5-1"></span>average pixel distance is employed as the metric for alignment. *Alignment Results*: In the alignment evaluation, we select scenes with well-defined textures, as illustrated in Fig. [3](#page-4-0) (b). After calculating the pixel distances, we observe that the average pixel error between the paired images is 0.2967 pixels. Throughout the entire time sequence, the pixel-level distance remains below 0.8 pixels, with the majority of errors being under  $0.5$  pixels, as exhibited in Fig. [3](#page-4-0) (c). These results demonstrate that the registration accuracy of our dataset reaches sub-pixel precision. *For further details, please refer to the appendix.*

# <span id="page-5-0"></span>5 METHODS

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**301 302 303 304 305 306 307** Overview: As shown in Fig. [4,](#page-5-1) our framework, SEE-Net, consists of four implementation parts: (a) Inputs and Preprocessing, (b) Encoding, (c) Decoding, and (d) Loss Function. The input is an image  $I_i$  and its corresponding events E. The output is a brightness-adjustable image  $I_o$ , where the brightness is controlled by the prompt  $B \in (0, 1)$ . During training, the brightness prompt B is calculated according to the target image. On the other hand, during testing,  $B$  can be freely set, with a default value of 0.5, which follows the exposure control constraint [\(Mertens et al.,](#page-11-12) [2009;](#page-11-12) [2007\)](#page-11-14). Overall, the SEE-Net  $f_{see}$  can be described by the Eq. [2](#page-5-2) to match our learning task in Sec. [3.](#page-2-0)

<span id="page-5-2"></span>
$$
I_o = f_{\rm see}(I_i, E, B). \tag{2}
$$

**309** Below, we elaborate the insights and implementation details of each part.

**310 311 312 313 314 315 316 317 318 319 320 321 322 323** Inputs and Processing: This part aims to transform initial inputs into features that retain original information for the encoding stage. The inputs consist of the image  $I_i$  and the events E, where  $I_i$  has a dimension of  $H \times W \times 3$  (with H and W representing the height and width, and 3 representing the color channel number). The event stream  $E$  is represented as a voxel grid [\(Tulyakov et al.,](#page-12-10) [2022\)](#page-12-10) with a dimension of  $H \times W \times M$ , where M represents the number of time slices of events. The events include color information [Scheerlinck et al.](#page-12-11) [\(2019\)](#page-12-11), which was overlooked in previous works, *e.g.*, [\(Liang et al.,](#page-11-4) [2024;](#page-11-4) [2023\)](#page-11-5). Specifically, this DVS346 sensor records events with Bayer Pattern [\(Lukac](#page-11-15) [et al.,](#page-11-15) [2005\)](#page-11-15). To effectively embed both the color and positional information during framework training, we design the position and bayer pattern embeddings, as shown in Fig. [4](#page-5-1) (a). The position and Bayer Pattern are denoted as a vector  $(x, y, bp)$ , where x, y is the pixel position, and by denotes the Bayer Pattern index, which takes a value from 0 to 3. We embed this vector into a higher-dimensional feature, termed as P, and concatenate it with the inputs. Two layers  $1 \times 1$  convolutions, denote  $f_e$ and  $f_i$ , are then applied to obtain the initial event features  $F_e$  and image features  $F_i$ . This process is described by the Eq. [3,](#page-5-3) where  $f_{cat}$  denotes the concatenation function.

<span id="page-5-3"></span>
$$
F_e = f_e(f_{cat}(E, P)), \quad F_i = f_i(f_{cat}(I_i, P)).
$$
\n(3)

<span id="page-6-4"></span>**324 325 326 327 328 329 330 331 332 333 334 335 336 337 338 Encoding:** In this stage, we aim to obtain the BLR by employing the event feature  $F_e$  to enhance the image feature  $F_i$ , facilitating noise reduction and the acquisition of broader light range information. Since  $F_e$  contains rich information about the lighting changes across different intensity levels, we use it as the source for representing the broader light range. However, event data only records changes in illumination, which differ fundamentally from the static RGB frame modality. This makes directly utilizing event data for broader light representation challenging. To address this, we employ a cross-attention [\(Liang et al.,](#page-11-16) [2021\)](#page-11-16) for feature fusion, producing the initial fused broad-spectrum feature  $F_1$ , expressed as  $F_1 = f_{c_0}(F_e, F_i)$ , where  $f_{c_0}$  is a cross-attention block. Then, inspired by previous works [\(Wang et al.,](#page-12-12) [2020\)](#page-12-12), we utilize sparse learning to generate residuals for  $F_1$  from the event features  $F_e$ . These residuals are progressively generated from the loop that executes L times. Multiple iterations are used because they allow the model to iteratively refine the residuals, capturing finer details and enhancing the feature representations by progressively integrating information from the events. A single loop of this process can be expressed as,  $F_{j+1} = f_l(F_e, F_1, F_j)$ , where  $f_l$  is a loop function that contains two cross-attention blocks as shown in Fig. [4](#page-5-1) (b), where  $F_i$  and  $F_{i+1}$ are the input and output of one loop. After L iterations, the final feature  $F<sub>L</sub>$  represents the BLR, as described by Eq. [4.](#page-6-1)

<span id="page-6-1"></span>
$$
F_L = f_{se}(F_e, F_1) = f_l(F_e, F_1, f_l(F_e, F_1, \ldots f_l(F_e, F_1, F_1))).
$$
\n(4)

**341 342 343 344 345 346 347 348 349 350 351 352 353 354 355 Decoding:** The objective of this part is to decode the BLR into a brightness-adjustable image  $I_0$ . In designing this decoder, we focus on two key insights: (1) The decoding process should be pixel-wise and efficient, allowing for greater flexibility during model deployment; (2) The embedding of the brightness information should be thorough and fully integrated. With these insights, we design the decoder with only a 5-layer MLP as shown in Fig  $4$  (c). Our decoder begins by encoding the brightness prompt  $B \in (0, 1)$  into an embedding vector. To effectively encode the high-frequency brightness prompt into features that are easier for the network to learn [\(Vaswani,](#page-12-13) [2017\)](#page-12-13), we introduce a learnable embedding, denoted as  $\mathbf{B} = f_{pe}(B) = f_{mlp}(f_{cat}(f_{mlp}(B), B))$ , which consists of two MLP layers. Through this embedding, the brightness prompt  $B$  is transformed into a vector  $B$ , matching the dimensions of the BLR channels. We then integrate this embedding  $\bm{B}$  into the decoder. To ensure the brightness prompt is fully incorporated and prevent information loss through multiple MLP layers, we employ a multi-step embedding approach, as displayed in Eq. [5,](#page-6-2) which guarantees that the brightness is progressively embedded throughout the decoding process. During the training phase, the prompt B is derived from the reference image by applying  $f<sub>b</sub>$  to calculate the global average brightness. In contrast, during the testing phase,  $B$  can be set freely, with a typical example being a value of 0.5.

<span id="page-6-2"></span>
$$
I_o = f_d(F_L, B) = f_{mlp}(B + f_{mlp}(B + ... f_{mlp}(B + F_L))).
$$
\n(5)

**357 358 359 360 361 362 363 364 365 Loss Function:** The purpose of our loss function is to supervise the prediction  $I_0$  using the ground truth  $I_t$ , with the corresponding brightness  $B = f_b(I_t)$ . The loss function consists of two main components: image reconstruction loss  $\mathcal{L}_i$  and gradient loss  $\mathcal{L}_q$ . The image reconstruction loss is Charbonnier loss [\(Lai et al.,](#page-11-17) [2018\)](#page-11-17), which effectively handles both small and large errors. Additionally, we employ gradient loss to improve the structural consistency of the output image. This is achieved by enforcing  $L_1$  constraints on the gradients of both the output and ground truth images. Therefore, the overall loss function is formulated as a weighted sum of the image loss and gradient loss, as exhibited in Eq. [6.](#page-6-3) Here,  $\nabla$  denotes the gradient operator, and  $\lambda_1$  and  $\lambda_2$  are the weights that balance the contributions of two loss terms.

<span id="page-6-3"></span>
$$
\mathcal{L}(I_o, I_t) = \lambda_1 \mathcal{L}_i + \lambda_2 \mathcal{L}_g = \lambda_1 \sqrt{(I_o - I_t)^2 + \epsilon^2} + \lambda_2 \|\nabla I_o - \nabla I_t\|.
$$
 (6)

<span id="page-6-0"></span>6 EXPERIMENTS

**339 340**

**356**

**370 371 372 373 374 375 376 377** Experimental Setting: *Implementation Details:* Our experiments use the Adam optimizer with an initial learning rate of  $2e - 4$  for all the experiments. We train our model for 40 epochs on the SDE dataset [\(Liang et al.,](#page-11-4) [2024\)](#page-11-4). On the SEE-600K dataset, we train for only 20 epochs, as SEE-600K is extremely large. All of our training is conducted on an HPC cluster, with a batch size of 2. To enhance data diversity, we apply random cropping to the images and perform random flips and rotations. *Evaluation Metrics:* We maintain consistency with previous methods [\(Liang et al.,](#page-11-4) [2024;](#page-11-4) [2023\)](#page-11-5) by using PSNR and SSIM [\(Wang et al.,](#page-12-14) [2004\)](#page-12-14). However, since our proposed new problem is highly challenging and most current approaches perform poorly on our SEE-600K dataset, we additionally introduce the  $L_1$  distance as a reference.

<span id="page-7-2"></span><span id="page-7-0"></span>Table 1: Comparison of different methods on the SDE dataset. The best performances is highlighted in **bold**. † refers to the original model for the HDR task, which is fine-tuned and trained on SDE





<span id="page-7-1"></span>

Figure 5: Visualization results on the SDE dataset.

**412 413 414 415 416 417 418 Dataset:** We conduct experiments on two real-world datasets: (1) SDE [\(Liang et al.,](#page-11-4) [2024\)](#page-11-4) comprises 91 scenes, with 76 for training and 15 for testing. Each scene includes a pair of low-light and normal-light images along with their corresponding events. (2) **SEE-600K** consists of 202 scenes, with each scene containing an average of four sets of videos under different lighting conditions, ranging from low light to bright light. During each training session, we randomly select one set of normal-light images as the reference and use the remaining sets as inputs. For example, for one scene with one low-light, two normal-light, and one high-light set, we generate six pairs of training data.

**419 420 421 422 423 424 425 426 427 428 429** Comparative Methods: We categorize the approaches we compare into four groups. Firstly, DCE [\(Guo et al.,](#page-10-14) [2020\)](#page-10-14) is a classical approach that can adjust the image brightness curve to achieve normal lighting. Secondly, there are strategies that only use images as input, including  $SNR$  ( $X<sub>u</sub>$ [et al.,](#page-13-6) [2022\)](#page-13-6), UFormer [\(Wang et al.,](#page-12-15) [2022\)](#page-12-15), LLFlow [\(Wu et al.,](#page-13-7) [2023\)](#page-13-7), and RetinexFormer [\(Cai et al.,](#page-10-7) [2023\)](#page-10-7). Thirdly, we consider methods that rely solely on events, *e.g.*, E2VID+ [\(Stoffregen et al.,](#page-12-1) [2020\)](#page-12-1). Tertiary, we examine event-guided low-light enhancement frameworks. This group includes single-frame input methods, *e.g.*, eSL-Net [\(Wang et al.,](#page-12-12) [2020\)](#page-12-12), [Liu et al.](#page-11-6) [\(2023\)](#page-11-6), [Wang et al.](#page-12-16) [\(2023a\)](#page-12-16) and EvLight [\(Liang et al.,](#page-11-4) [2024\)](#page-11-4), as well as multi-frame input strategies like EvLowLight [\(Liang et al.,](#page-11-5) [2023\)](#page-11-5). Furthermore, we also compared the HDR reconstruction method HDRev[Yang et al.](#page-13-1) [\(2023\)](#page-13-1). We retrain all methods, following the open-source code when available; for approaches without open-source code, we replicate them based on their respective papers.

**430 431** Comparative on SDE Dataset: The results from our comparative experiments, shown in Tab. [1,](#page-7-0) reveal several key insights: (1) performance limitations of single-modal methods: Methods utilizing only one modality exhibit limited performance, as shown in Tab. [1.](#page-7-0) This trend underscores the

# **436 456 Training<br>Dataset** Exaining Methods **low light** high light high light normal light normal light normal light  $\frac{1}{PSNP} = \frac{1}{SNP} = \frac{$ PSNR SSIM  $L_1$  PSNR SSIM  $L_1$  PSNR SSIM  $L_1$ SDE DCE [\(Guo et al.,](#page-10-14) [2020\)](#page-10-14) 9.10 0.0968 0.3572 6.26 0.3419 0.4649 10.79 0.3992 0.2524<br>eSL Net (Wang et al., 2020) 11.92 0.3275 0.2703 6.66 0.1672 0.4001 7.65 0.2685 0.3481 eSL Net [\(Wang et al.,](#page-12-12) [2020\)](#page-12-12) 11.92 0.3275 0.2703 6.66 0.1672 0.4001 7.65 0.2685 0.3481<br>Liu et al. (2023) 12.41 0.4001 0.2487 5.53 0.1950 0.4534 6.58 0.2805 0.4129 [Liu et al.](#page-11-6) [\(2023\)](#page-11-6) 12.41 0.4001 0.2487 5.53 0.1950 0.4534 6.58 0.2805 0.4129<br>EvLowLight (Liang et al., 2023) 12.68 0.4341 0.2338 4.11 0.3071 0.6062 7.01 0.3950 0.4520 EvLowLight [\(Liang et al.,](#page-11-5) [2023\)](#page-11-5) 12.68 0.4341 0.2338 4.11 0.3071 0.6062 7.01 0.3950 0.4520 EvLight [\(Liang et al.,](#page-11-4) [2024\)](#page-11-4) 13.07 0.4651 0.2337 5.12 0.1005 0.4842 6.29 0.2805 0.4336<br>SEENet **14.84 0.5693 0.1779** 3.84 0.2119 0.6123 5.36 0.2980 0.5056 SEENet 14.84 0.5693 0.1779 3.84 0.2119 0.6123 5.36 0.2980 0.5056 **SEE** eSL Net [\(Wang et al.,](#page-12-12) [2020\)](#page-12-12) 11.95 0.3845 0.2421 12.84 0.4660 0.2076 13.45 0.5682 0.1957<br>
EvLowLight + (Liang et al. 2023) 12.83 0.4511 0.2151 12.79 0.4696 0.2084 13.04 0.5531 0.2144 *EvLowLight* † [\(Liang et al.,](#page-11-5) [2023\)](#page-11-5) 12.83 0.4511 0.2151 12.79 0.4696 0.2084 13.04 0.5531 0.2144 [Liu et al.](#page-11-6) [\(2023\)](#page-11-6) 13.48 0.5068 0.1946 12.30 0.4766 0.2221 13.70 0.5474 0.2151 EvLight [\(Liang et al.,](#page-11-4) [2024\)](#page-11-4) 13.70 0.5150 0.1960 13.45 0.4918 0.1990 13.63 0.5924 0.2004<br>SEENet 18.77 0.6303 0.0971 19.21 0.6675 0.0806 20.92 0.8002 0.0606 SEENet 18.77 0.6303 0.0971 19.21 0.6675 0.0806 20.92 0.8002 0.0606 (a) Events (b) Inputs Image (Low-light) (c) Normal-light Image (d) SEE Net (Ours) 27.17/0.7761 (e) Liu et. al. 15.98/0.6205 (f) eSL Net 14.32/0.5304 (g) EvLight 26.65/0.7145 (h) SEE Net (Ours) Prompt 0.5 (i) Irvents (j) Inputs Image (Low-light) (k) Normal-light Image (l) SEE Net (Ours) 26.85/0.8937 (m) Liu et. al. 18.90/0.6032 (n) eSL Net 17.30/0.5386 (o) EvLight 16.63/0.4684 (p) SEE Net (Ours) Prompt 0.5

<span id="page-8-2"></span><span id="page-8-0"></span>Table 2: Evaluation on the SEE-600K dataset, with methods trained on both the SDE and SEE-600k. *EvLowLight* † *refers to this method trained after downsampling the dataset by 10 times.*

<span id="page-8-1"></span>Figure 6: Visual examples of low-light enhancement and high-light recovery on the SEE-600K dataset.

**469 470 471 472 473 474 475 476 477** necessity of integrating both modalities for enhanced results, as shown in Fig.  $5$  (f). (2) effectiveness of event-guided methods: In contrast, event-guided image methods demonstrate significantly better performance. These approaches leverage the complementary strengths of both events and traditional images, leading to better outcomes in low-light conditions, as shown in Fig.  $5(g-i)$  $5(g-i)$ . (3) impact of indoor and outdoor conditions: Notably, performance in low-light indoor scenarios is inferior to that in outdoor settings, as shown in Fig.  $5$  (e). This discrepancy may be attributed to the issues of flickering light sources commonly found indoors [\(Xu et al.,](#page-13-8) [2023a\)](#page-13-8). Our SEE-Net consistently achieves the best results across both scenarios, with a model size of just  $1.9M-10\%$  parameter count of other SOTA methods—demonstrating its efficiency and compactness in low-light image enhancement.

**478 479 480 481 482 483 484 485** Comparative on SEE-600K Dataset: The results presented in Tab. [2](#page-8-0) illustrate the performance of various methods across different lighting conditions on the SEE-600K dataset. (1) trained on SDE: Models trained on the SDE dataset maintain a reasonable level of performance when tested on the SEE-600K dataset, particularly in low-light conditions. Notably, the DCE [Guo et al.](#page-10-14) [\(2020\)](#page-10-14) achieves the best results in high-light scenarios, underscoring its excellent generalization capabilities for its self-supervised approach. (2) trained on SEE-600K: Models trained on the SEE-600K dataset exhibit improved performance in both low-light and high-light conditions. Our proposed SEE-Net method stands out as the best performer, as shown in Tab[.2](#page-8-0) and Fig. [6.](#page-8-1) This achievement is due to our innovative use of prompt adjustments, which effectively resolve the ambiguity often seen in

			. r					
Case	<b>Bayer</b> Pattern	Encoding	Loop	Prompt <b>Embedding</b>	Cascade	Prompt Merge	<b>PSNR</b>	<b>SSIM</b>
	$f_{pe}$	$f_{ca}$	20	$f_{pe}$		$^{+}$	23.57	0.7724
2		$_{,a}$	20	$f_{pe}$		$^{+}$	22.94	0.7686
3	$f_{pe}$	$add + conv$	20	$f_{pe}$		$^{+}$	22.40	0.7224
	$ _{pe}$	$cat + conv$	20	$_{\it Ipe}$			22.84	0.7298
5	$f_{pe}$	$f_{ca}$	10	$f_{pe}$		$^{+}$	22.18	0.6812
6	$f_{pe}$	$f_{ca}$	20	sin		$^{+}$	23.08	0.7692
	$\int pe$	$_{Ica}$	20	$f_{pe}$		$^{+}$	22.26	0.7713
8	$_{\rm Jpe}$	$f_{ca}$	20	$f_{pe}$		$\times$	22.94	0.7893
	 <b>WARRIERS</b>		  <b>ARK 88</b> <b></b>	<b><i><u><u>ELES TETT</u></u></i></b> ------- ------  ------- <b><i><u>AAAAAAA</u></i></b> <b><i><u>AAAAAAA</u></i></b> <b>BARN SERIE</b>	<b>BARN BRACK LEARN BRACK LEARN BRACK LEARN BRA</b>		. .	
	  <b>ARREL 8 8 8</b> <b>Free 848</b>		$\frac{1}{2}$ .  . <b></b> ----- <b></b> <b>STATISTICS</b> <b>CONTRACTOR</b>	. <b></b> ------ <b>CERTIFICATE</b> <b>ALAN XXX</b> <b>ALAN YEAR</b> <b>ALAN YEAR</b> <b>WEEK ARES</b> <b><i><u>APPEND</u></i></b> <b>TEXAS &amp; RAS</b> <b><i><u>ARK AVA</u></i></b> <b><i><u>ARRA EVE</u></i></b> <b><i><u>ARRE LEX</u></i></b> <b><i><u>ARKA 64.54</u></i></b> <b>ALL A V 1999</b> <b>XXXXXXXX</b> <b>NAME OF GROOM</b> <b>STATISTICS</b> <b>THEF READ</b> <b>ANDREWS</b>	<b>CONTRACTORS CONTRACTORS CONTRACTORS</b> <b>SHER STATISTICS AND LOCAL AND</b> <b>CERTIFICATE CERTIFICATE CERTIFICATE CONTRACTOR</b> a bond in a 4 days are not the agreement of a property and		.  . 1.1.1.1	
	<b>HARR KAR</b> <b>STAR REA</b> $-100$		<b></b> <b>STATISTICS</b>	<b>A R A R &amp; R W</b> <b>AMAX AND</b> <b>AREE AVES</b> <b>CERE CALL</b> <b>STEE RASH</b> <b>STEED CALL</b> <b>BASE VIX</b> <b>MARK VAN</b> <b>MARK VANS</b>	ANDE ANNE <sub>BANDE</sub> ANNI <sub>BANDE</sub> ANNI <sub>BANDE</sub> ANN <b>CARS RASS LANGE RASS</b> VALUE OF THE			
		$0.3^+$			$\overline{0.7}$			
	(a) Input Events & Image			(b) Prediction with Brightness Prompt $B$ from 0.3 to 0.7				(c) Normal-light

<span id="page-9-2"></span><span id="page-9-0"></span>Table 3: Ablation studies.  $f_c$  indicates cross-attention.  $f_{pe}$  stands for learning-based embedding.

<span id="page-9-1"></span>Figure 7: Visualization of brightness adjustment using varying brightness prompts B from 0.3 to 0.7, showing smooth brightness transitions in SEE-600K dataset. For more visualizations, see the Appendix.

 enhancement processes. Overall, these results highlight the effectiveness of our approach across diverse lighting conditions, further validating its robustness. (3) advantages of prompt adjustments: Unlike previous methods, Fig. [6,](#page-8-1) that are limited to one-way mapping, our approach with prompt adjustments demonstrates significant advantages, as shown in Fig. (h,p). Prompt adjustments allow us to produce image quality that surpasses the ground truth, Fig. (d,i), regardless of whether low-light or high-light conditions are used as input. When the prompt is set to 0.5, the output achieves optimal brightness and sharp textures. *For additional visualization, please refer to the appendix.*

 Ablation and Analytical Studies: In this ablation study (Tab[.3\)](#page-9-0), we analyze the impact of various components using Case #1 as the baseline. (1) bayer pattern embedding: Removing the bayerpattern embedding (Case #2) leads to a performance drop, indicating it enhances accuracy but is not the most critical factor. (2) encoding: Replacing the cross-attention module  $f_c$  with a convolutional layer in both Case #3 (add) and Case #4 (concat) leads to significant performance degradation, underscoring the critical role of cross-attention. (3) loop iterations: Reducing loop iterations from 20 to 10 (Case #4) causes a performance decline, indicating sufficient iterations are necessary for refinement. (4) prompt embedding: Switching the prompt embedding from  $f_{pe}$  to a sine function [\(Vaswani,](#page-12-13) [2017\)](#page-12-13) (Case  $#5$ ) yields similar performance but doesn't surpass the learned embedding. (5) prompt merge: Disabling prompt merge (Case #6) results in a slight performance drop, indicating its importance for optimal results. (6) multi-prompt adjustment: Fig[.7](#page-9-1) shows the output under multiple prompts. The input consists of a low-light image and events. When using gamma correction to brighten the low-light image, significant noise is introduced (Fig[.7](#page-9-1) (a)). However, our outputs with varying prompts effectively control brightness while reducing noise (Fig[.7](#page-9-1) (b)), demonstrating the flexibility and robustness of our method in post-processing. *Due to space limitations, please refer to the appendix for more information.*

 

# 7 CONCLUSION

 In this paper, we proposed a new research problem: how to use events to adjust the brightness of images across a wide range of lighting conditions, from low light to high light. To address this challenge, we made the following contributions. (1), we developed a physical model and formally

 defined the problem of brightness adjustment using events, providing a solid theoretical foundation. (2), we introduced a new spatiotemporal registration algorithm based on a robotic arm and collected a large-scale dataset, SEE-600K, to overcome alignment issues and support our research. (3), we presented SEE-Net, a novel and compact framework capable of accepting input images with a wide range of illumination and producing output images with adjustable brightness. (4), we conducted extensive experiments to demonstrate the effectiveness of our method.

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# Appendix

**768 769 770** To address the reviewers' feedback, we have added the following three sections to the supplementary material:

**771 772 773** (1) More Details About the DVS346 Sensor: We provide additional information on the sensor's specifications, particularly regarding noise and image quality, to contextualize the limitations of the APS frames in our dataset.

**774 775 776** (2) Differences Between Brightness Adjustment and HDR Reconstruction: We clarify the differences between our brightness adjustment task and HDR reconstruction, focusing on objectives, challenges, and data construction methods, supported by mathematical formulations.

**777 778 779 780 781** (3) Output Visualizations of Different Prompts: We include visual examples showing how our network processes inputs under extreme low-light and high-light conditions using various brightness prompts, directly addressing how the brightness prompt  $B$  influences the outputs from different input images.

In the final paper, we will organize the Appendix accordingly. For now, we have placed these sections at the beginning of the supplementary material for the reviewers' convenience.

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# A MORE DETAILS ABOUT THE DVS346 SENSOR

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**787 788 789 790 791** In our experiments, we employed the DVS346 event camera, a sensor capable of simultaneously outputting asynchronous events and synchronous image frames (APS frames). Despite its widespread use in the academic community, the DVS346 has inherent limitations that affect the quality of the captured images, particularly due to various noise factors. Understanding these parameters is crucial for contextualizing the performance of our proposed methods.

**792 793 794** The specifications of the DVS346 sensor are detailed in Tables [4](#page-14-1) and [5.](#page-15-1) Below, we explain the significance of each parameter, emphasizing those related to noise, to illustrate the image quality from this sensor.

**795 796 797 798 799 800 801 802 803 804 805 806 807 808** Events: *Spatial resolution:* refers to the number of pixels in the sensor array, which in this case is 346  $\times$  260 pixels. *Temporal resolution:* of 1,  $\mu$ s indicates the sensor's ability to detect rapid changes in brightness, allowing for precise temporal event detection. This high temporal resolution is advantageous for capturing fast-moving scenes. *Maximum throughput:* of 12 MEPS means the sensor can handle up to 12 million events per second, which is essential for recording scenes with a lot of motion without losing data. *Typical latency:* of less than 1, ms ensures minimal delay between the occurrence of an event and its registration by the sensor, which is important for real-time applications. *Dynamic range:* of approximately 120, dB allows the event sensor to operate effectively under a wide range of lighting conditions, from very dark to very bright environments. This high dynamic range is a key advantage of event-based cameras. *Contrast sensitivity*: represents the minimum percentage change in brightness required to generate an event. The sensor has a contrast sensitivity of 14.3% for ON events and 22.5% for OFF events. While higher contrast sensitivity reduces noise by preventing the sensor from triggering on minor fluctuations, it may also cause it to miss subtle changes in brightness.

**809 Frame:** *Spatial resolution:* for the APS frames is the same as the event output, limiting the detail in the captured images. Frame rate: of 40, FPS indicates that the sensor captures 40 frames per second.

<span id="page-15-1"></span><span id="page-15-0"></span>

<span id="page-15-2"></span>Figure 8: Examples of other datasets from the event-based vision community [\(Scheerlinck et al.,](#page-12-11) [2019;](#page-12-11) [Cui](#page-10-4) [et al.,](#page-10-4) [2024;](#page-10-4) [Wang et al.,](#page-12-17) [2023b;](#page-12-17) [Liang et al.,](#page-11-4) [2024\)](#page-11-4). Although the DVS 346 camera suffers from insufficient dynamic range and noise, it is still the data acquisition device that can best support the training of various event vision tasks at this stage.

**835 836 837 838 839 840 841 842 843 844** *Dynamic range:* of 55, dB is significantly lower than that of the event output. This limited dynamic range means the APS frames struggle with scenes that have both very bright and very dark areas, leading to overexposure or underexposure in parts of the image. *Fixed Pattern Noise (FPN):* of 4.2% refers to the non-uniformity in pixel responses, where each pixel may have a slightly different baseline level of response due to manufacturing inconsistencies. High FPN manifests as a static noise pattern over the image, degrading the visual quality. *Dark signal*: of 18,000, e<sup>−</sup>/s represents the amount of charge accumulated by a pixel in the absence of light. A high dark signal increases the baseline noise level, especially noticeable in low-light conditions, resulting in grainy images. *Readout noise:* of 55, e<sup>−</sup> is the noise introduced during the process of reading the pixel values from the sensor. This electronic noise adds uncertainty to the pixel values, further reducing image clarity and detail, particularly in darker regions where the signal level is low.

**845 846 847 848 849 850 851 852 853 854 855** Impact on Image Quality: The combination of these parameters adversely affects the image quality of the APS frames produced by the DVS346 sensor: (1) A dynamic range of  $55 \text{ dB}$  is insufficient for high-contrast scenes, causing loss of detail in shadows (underexposure) or highlights (overexposure). This limitation means that the APS frames cannot effectively capture scenes with both bright and dark regions simultaneously. (2) High levels of **Fixed Pattern Noise** introduce consistent noise patterns across the image, which are difficult to remove and can be distracting in the final output. (3) The significant **dark signal** contributes to increased noise, especially in low-light conditions where the actual signal from the scene is weak. This results in a lower signal-to-noise ratio (SNR), making the images appear grainy or speckled. (4) Elevated readout noise further degrades image quality by adding random variations to the pixel values during the readout process, obscuring fine details and reducing overall sharpness.

**856 857 858 859 860 861 862** These noise-related issues collectively lead to suboptimal image quality in the APS frames, with noticeable artifacts such as blurriness, graininess, and loss of detail. Understanding the limitations of the DVS346 sensor is essential for interpreting the results of our research. While the sensor's APS frames have quality constraints due to noise and limited dynamic range, the event output excels in capturing high temporal resolution and wide dynamic range changes. Our work leverages the strengths of the event data to adjust image brightness across various lighting conditions, mitigating some of the APS frame limitations.

**863** Despite the challenges posed by the sensor's noise characteristics, the DVS346 remains a valuable tool in event-based vision research [\(Scheerlinck et al.,](#page-12-11) [2019;](#page-12-11) [Cui et al.,](#page-10-4) [2024;](#page-10-4) [Wang et al.,](#page-12-17) [2023b;](#page-12-17) [Liang](#page-11-4)

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**864 865 866 867 868** [et al.,](#page-11-4) [2024\)](#page-11-4) due to its accessibility and the richness of the event data it provides, as shown in Fig. [8.](#page-15-2) As technology advances, we anticipate that future sensors will offer improved image quality with reduced noise levels, enhancing the potential for high-quality event-based imaging. In the meantime, acknowledging and addressing these limitations allows us to develop algorithms that compensate for the sensor's shortcomings, contributing to the advancement of event-based vision applications.

# B DIFFERENCES BETWEEN BRIGHTNESS ADJUSTMENT AND HDR **RECONSTRUCTION**

In this section, we discuss the fundamental differences between our proposed brightness adjustment task using event cameras and the traditional High Dynamic Range (HDR) reconstruction task. We highlight the distinctions in objectives, challenges, and data construction methodologies, supported by mathematical formulations for clarity.

#### **878** Different Objectives:

**879 880 881 882** The primary goal of HDR reconstruction is to expand the dynamic range of an image, capturing details in both dark and bright regions that exceed the capability of standard Low Dynamic Range (LDR) sensors. Mathematically, HDR imaging seeks to recover a radiance map  $R(x)$  that represents the true scene radiance over a wide dynamic range:

**883 884 885**

$$
R(x) = f^{-1}(I_{\text{LDR}}(x)),
$$
\n(7)

 $\hat{I_{\text{rgb}}} = f_{\text{see}}(I_{\text{rob}}, E; B),$ (8)

where  $I_{\text{LDR}}(x)$  is the observed LDR image, and  $f^{-1}$  is the inverse of the camera response function.

In contrast, our brightness adjustment task focuses on modifying the exposure level of an image to enhance visibility and recover lost details due to underexposure or overexposure, without necessarily expanding the dynamic range. The objective is to obtain an adjusted image  $I_{\text{rgb}}$  from an input image  $I_{rgb}$  and event data  $E(x, t)$ :

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where  $f_{\text{see}}$  is our proposed adjustment function,  $E$  represents the event stream, and  $B$  is the brightness prompt controlling the desired exposure level.

#### **897** Different Challenges:

HDR reconstruction faces the challenge of accurately merging multiple images captured at different exposure levels to create a single image with an expanded dynamic range. This often requires precise alignment and handling of motion between exposures to avoid ghosting artifacts. The mathematical formulation involves combining N images  $\{I_i(x)\}_{i=1}^N$  with corresponding exposure times  $\{t_i\}_{i=1}^N$ :

$$
R(x) = \frac{\sum_{i=1}^{N} w(I_i(x)) \cdot f^{-1}(I_i(x))}{\sum_{i=1}^{N} w(I_i(x))},
$$
\n(9)

**906 907** where  $w(I_i(x))$  is a weighting function that emphasizes well-exposed pixels.

**908 909 910 911 912 913** Our brightness adjustment task, on the other hand, deals with the challenge of adjusting images captured under various lighting conditions using the high temporal resolution and dynamic range of event data. Unlike HDR reconstruction, we do not require multiple images at different exposures. Instead, we leverage events to infer illumination changes and guide the brightness adjustment of a single input image. The adjustment function  $f_{see}$  must effectively fuse spatial image data and temporal event information:

**914 915**

**916**

$$
I_{\text{adj}}(x) = f_d \left( f_{se}(I_{rgb}, E), B \right),\tag{10}
$$

**917** where  $f_{se}$  is an encoder that extracts features from the input image and events, and  $f_d$  is a decoder that generates the adjusted image based on the brightness prompt B.



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<span id="page-17-0"></span>Figure 9: Under the same scene, with low-light and high-light images as inputs, we compare the outputs generated using a series of prompts. The inputs are the original image, events, and the visualization of the original image after gamma correction. Almost all the contours and details in the original image are lost.

 

### **972 973** Different Data Construction Methods:

**974 975 976** Constructing datasets for HDR reconstruction typically involves capturing multiple images of the same scene at different exposure levels, requiring static scenes or sophisticated alignment techniques to handle motion. The ground truth HDR image is often synthesized by merging these exposures.

**977** Mathematically, for each scene, we collect  $N$  images:

$$
\{I_i(x)\}_{i=1}^N, \quad \text{with exposure times } t_1 < t_2 < \cdots < t_N,\tag{11}
$$

**981** and compute the ground truth radiance  $R(x)$  as shown earlier.

**982 983 984 985** For our brightness adjustment task, data construction is more straightforward and scalable. We capture pairs of images and corresponding event data under varying lighting conditions using Neutral Density (ND) filters to simulate different exposures. Each scene provides synchronized data without the need for multiple exposure times or complex alignment:

$$
(I_{rgb}, E, I_{rgb}^{\hat{}}), \qquad (12)
$$

**989 990 991** where  $I_{\text{rgb}}$  is the ground truth image at the desired exposure. The use of events allows us to handle dynamic scenes effectively, as the high temporal resolution of events captures rapid changes in illumination.

**992 993 994 995 996** In essence, while HDR reconstruction aims to create images with an expanded dynamic range by combining multiple exposures, our brightness adjustment task seeks to adjust the exposure of images using event data to recover lost details without extending the dynamic range. Our approach is more practical for real-world applications where capturing multiple exposures is impractical or impossible.

**997 998 999 1000** By formulating the problem differently and leveraging the unique properties of event cameras, we address challenges specific to brightness adjustment under diverse lighting conditions. This includes handling dynamic scenes and providing fine-grained control over image brightness through prompts.

**1001 1002 1003** Our dataset construction method is also more scalable, enabling us to create a large dataset without the complexities involved in HDR dataset creation. This allows for training more robust models suited to real-world scenarios.

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#### **1005** C OUTPUT VISUALIZATIONS OF DIFFERENT PROMPTS

**1007 1008 1009 1010** The Fig. [9](#page-17-0) demonstrates the effectiveness of our network in reconstructing images across a broad range of lighting conditions using events. We input both extremely low-light and overexposed images—where the original contours and details are significantly degraded or lost—into our network to observe how it handles varying input brightness levels when the same brightness prompt is applied.

**1011 1012 1013 1014** Our network leverages the high dynamic range and temporal resolution of events to recover lost details in both underexposed and overexposed scenarios. By integrating events, which captures pixel-level changes in brightness over time, the network compensates for the deficiencies of the input images regardless of their initial exposure levels.

**1015 1016 1017 1018 1019** We present the results corresponding to brightness prompts ranging from 0.2 to 0.8, allowing for fine-grained control over the brightness of the output images. Each prompt value is applied to both the extremely low-light and overexposed input images. Despite the drastic differences in the original brightness of the inputs, the outputs generated with the same brightness prompt are remarkably consistent in terms of exposure and detail.

**1020 1021 1022 1023 1024 1025** This observation directly answers the reviewer's question: when reconstructing a bright image (*e.g*., setting  $B = 0.8$ ) from two different input images—one dark and one bright—the network produces output images that are both well-exposed and visually similar. Although the low light input image produced some artifacts. This demonstrates that the output is primarily determined by the brightness prompt  $B$ , rather than the original brightness of the input images. The network effectively adjusts the input images to the desired brightness level specified by the prompt, utilizing the event data to recover or suppress details as needed.



### <span id="page-20-0"></span>**1080 1081** D MORE DETAILS FOR RESEARCH PROBLEM DEFINITION

**1082 1083 1084 1085 1086** Imaging is the process of capturing light from a scene, which can be represented as a radiance field  $L(t)$  that varies over time t. The intensity of ambient light in real-world environments spans a wide range, from approximately 0.1 lux in low-light conditions to over 1e6 lux under bright sunlight. The goal of our learning task is to accurately recover  $L(t)$  and transform it into a visual representation that is suitable for human perception.

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# **1088** Sensor Signal Acquisition and Noise Modeling:

**1089 1090 1091 1092** Cameras equipped with active pixel sensors record light signals through an exposure process. During the exposure time  $t_e$ , the sensor integrates incoming photons to produce a voltage  $V$ . The number of photons  $k$  detected is a random variable following a Poisson distribution due to the quantum nature of light:

$$
k \sim \mathcal{P}(\lambda), \quad \lambda = \eta \int_{t_e} L(t)dt,
$$
\n(13)

**1095** where:

- $\lambda$  is the expected number of photons,
- $\eta$  is the quantum efficiency of the sensor,
- $L$  is the light intensity,
- $t_e$  is the exposure time.

**1103 1104** The voltage  $V$  generated by the sensor is proportional to the number of detected photons and is given by:

$$
V = Gk + N_d,\tag{14}
$$

**1106** where:

**1107 1108 1109**

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**1105**

•  $G$  is the sensor gain, usually a circuit amplifier, •  $N_d \sim \mathcal{N}(\mu_d, \sigma_d^2)$  represents the dark current noise, typically modeled as Gaussian noise

with mean  $\mu_d$  and variance  $\sigma_d^2$ .

- **1112** The RAW image intensity  $I_{\text{raw}}$  is obtained by quantize the voltage  $V$ :
- **1113 1114**

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**1131 1132 1133**

# $I_{\text{raw}} = \mathcal{Q}(V) = \mathcal{Q}(Gk + N_d),$  (15)

**1115 1116 1117** where  $Q$  is the quantization function converting continuous voltage signals into discrete digital values, typically ranging from 8 *bits* to 14 *bits*.

### **1118** Image Signal Processing (ISP)

**1119 1120 1121** The RAW image  $I_{\text{raw}}$  undergoes an image signal processing pipeline  $f_{\text{isp}}$  that includes steps such as denoising [\(Buades et al.,](#page-10-16) [2005\)](#page-10-16), demosaicing [\(Li et al.,](#page-11-18) [2008\)](#page-11-18), color correction [\(Gasparini & Schettini,](#page-10-17) [2003\)](#page-10-17), and tone mapping [\(Debevec & Gibson,](#page-10-18) [2002\)](#page-10-18) to produce the final RGB image:

$$
I_{\rm rgb} = f_{\rm iso}(I_{\rm raw}).\tag{16}
$$

# **1124 1125** Characteristics of Accurate Exposure

**1126** An accurate exposure process aims to produce  $I_{\text{rgb}}$  with the following characteristics:

**1127 1128 1129 1130** 1. Accurate Exposure: The mean pixel intensity of  $I_{\text{rgb}}$  falls within a desirable range for human observation, typically normalized between 0.4 and 0.7 [\(Mertens et al.,](#page-11-12) [2009\)](#page-11-12):

$$
0.4 \le \frac{1}{N} \sum_{i=1}^{N} I_{\text{rgb}}^{(i)} \le 0.7,\tag{17}
$$

where  $N$  is the total number of pixels.

<span id="page-21-0"></span>**1134 1135 1136 1137 1138 1139 1140 1141 1142 1143 1144 1145 1146 1147 1148 1149 1150 1151 1152 1153 1154 1155 1156 1157 1158 1159 1160 1161 1162 1163 1164 1165 1166 1167 1168 1169 1170 1171 1172 1173 1174 1175 1176 1177 1178 1179 1180 1181 1182 1183 1184 1185** 2. Noise-Free: The influences of dark current noise  $N_d$  and photon shot noise  $N_s$  are minimized or eliminated:  $\text{Var}(I_{\text{rgb}}) \approx \text{Var}(G\eta \int_{t_e})$  $L(t)dt$ , (18) implying that the variance due to noise is negligible. 3. Color Neutrality: The image has no color cast; the grayscale values computed from each RGB channel are approximately equal [\(Buchsbaum,](#page-10-11) [1980\)](#page-10-11):  $f_{\text{gray}}(I_r) \approx f_{\text{gray}}(I_q) \approx f_{\text{gray}}(I_b),$  (19) where  $I_r$ ,  $I_q$ , and  $I_b$  are the red, green, and blue channels of  $I_{rgb}$ , and  $f_{gray}$  is a function mapping RGB values to grayscale. Limitations of Traditional Cameras Traditional cameras have a limited dynamic range of approximately  $80dB$ , which often results in loss of detail in scenes with high contrast. Under extreme lighting conditions, images may exhibit overexposed highlights or underexposed shadows, leading to insufficient edge and texture information. Advantages of Event Cameras Event cameras overcome these limitations by offering: • High Dynamic Range: Greater than 120 dB, allowing them to handle extreme lighting variations. • High Temporal Resolution: Less than 1 ms, enabling them to capture fast-changing scenes. Event cameras operate asynchronously by detecting changes in illumination at each pixel. The output is a stream of events, each represented as:  $(x, y, t, p),$  (20) where: •  $(x, y)$  are the pixel coordinates, •  $t$  is the timestamp, •  $p \in \{+1, -1\}$  indicates the polarity (increase or decrease in light intensity). Event Generation Mechanism An event is generated at a pixel  $(x, y)$  when the change in the logarithm of the light intensity exceeds a predefined threshold  $C$ :  $\Delta L(x, y, t) = \log(L(x, y, t)) - \log(L(x, y, t_k)) = pC,$  (21) where: •  $L(x, y, t)$  is the light intensity at time t, •  $t_k$  is the timestamp of the last event at pixel  $(x, y)$ , •  $p$  is the polarity, •  $C$  is the contrast sensitivity threshold. This condition can also be expressed in terms of relative intensity change:  $L(x, y, t)$  $\frac{L(x, y, t)}{L(x, y, t_k)} = e$  $p^C$ . (22) Proposed Model for Illumination Recovery Given the high dynamic range and temporal resolution of event cameras, we aim to utilize an images

**1186 1187**  $I_{\text{rgb}}$  and corresponding events E to recover the scene's illumination  $L(t)$  and present it in a humanfriendly format. However, due to the extensive theoretical range of  $L(t)$ , we introduce a brightness control prompt  $B$  to adjust the output image's mean brightness.



 where:

- $f_{\text{see}}$  is a function designed to enhance the input image  $I_{\text{reb}}$  using the events E and adjust the brightness according to  $B$ ,
- $\hat{I}_{\text{rgb}}$  is the output image with improved exposure,
- B is a user-defined parameter representing the desired mean brightness of  $I_{\text{rgb}}$ :

<span id="page-22-0"></span>
$$
B = \frac{1}{N} \sum_{i=1}^{N} I_{\text{rgb}}^{\hat{i}i}
$$
 (24)

# Benefits of the Proposed Approach

- 1. **Robust Training**: By presetting the parameter  $B$  during the training phase, the model can mitigate biases present in the training dataset, leading to more generalized performance.
- 2. Flexibility in Usage: During inference, setting  $B = 0.5$  (assuming pixel values are normalized between 0 and 1) aligns with common exposure levels, but users can adjust  $B$  for creative control over the image's brightness and exposure, enabling image adjustments and editing capabilities.

# E TEMPORAL IMU REGISTRATION ALGORITHM

 In this section, we provide a more detailed description of our IMU data registration algorithm, which aligns a source sequence  $S$  and a target sequence  $T$  by finding the optimal temporal bias  $b$  and matching length  $l$  that minimize the distance between them. Due to the high sampling rate of IMU data (1000 *Hz*), an exhaustive search over all possible biases is computationally prohibitive. Therefore, we introduce a multi-level iterative strategy that efficiently approximates the optimal alignment.

### IMU Data Calibration and Stability

 Fig. [10](#page-22-0) illustrates the calibration results of our IMU sensor over a one-hour period during which the sensor remained stationary. From this figure, we observe that the IMU's measurement errors are stable over long durations and do not increase over time. The deviations in the accelerometer's three axes and the gyroscope's three axes are consistent, indicating reliable sensor performance. Through calibration,

	ັ				
1245	<b>Sensor</b>	Axis	<b>Bias</b>	Variance	<b>Standard Deviation</b>
1246	Accelerometer	Х	$-0.009256$	$5.836 \times 10^{-6}$	0.002416
1247	Accelerometer		0.993344	$6.196 \times 10^{-6}$	0.002489
1248	Accelerometer	7	$-0.048622$	$1.348 \times 10^{-5}$	0.003672
1249	Gyroscope	Х	1.081781	0.010550	0.102711
1250	Gyroscope		$-1.791223$	0.011102	0.105365
1251	Gyroscope	Z	$-0.697237$	0.011360	0.106582

**1243 1244** Table 6: Calibration results showing biases, variances, and standard deviations for each axis of the accelerometer and gyroscope.

<span id="page-23-1"></span><span id="page-23-0"></span>**1242**

**1253 1254 1255 1256 1257 1258** we corrected these biases during preprocessing to enhance measurement accuracy. Specifically, for the camera used in our dataset collection, the calibrated IMU errors are quantified shown in Tab. [6.](#page-23-1) These low variance values indicate that the IMU's measurement noise is within an acceptable and small range, affirming that our calibration process effectively corrects sensor deviations. Consequently, we can achieve accurate results in our data registration by leveraging the stability of the IMU sensor. The specific implementation steps of our calibration process are detailed below.

### **1259** IMU Data Preprocessing with Kalman Filter

**1260 1261 1262** We first denoise the raw IMU data using a Kalman filter [Mirzaei & Roumeliotis](#page-12-8) [\(2008\)](#page-12-8). For each IMU sequence (source and target), we model the system as:

**1263 1264 1265**

**1266**

$$
\mathbf{x}_k = \mathbf{F} \mathbf{x}_{k-1} + \mathbf{w}_{k-1},\tag{25}
$$

$$
\mathbf{z}_k = \mathbf{H}\mathbf{x}_k + \mathbf{v}_k,\tag{26}
$$

**1267 1268** where  $\mathbf{x}_k \in \mathbb{R}^6$  is the state vector at time k, consisting of accelerometer and gyroscope measurements:

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\n
$$
\mathbf{x}_k = \begin{bmatrix} \mathrm{acc}_x \\ \mathrm{acc}_y \\ \mathrm{gyr}_x \\ \mathrm{gyr}_y \\ \mathrm{gyr}_z \end{bmatrix}_k
$$

**1276 1277 1278**  $\mathbf{F} \in \mathbb{R}^{6 \times 6}$  is the state transition matrix (identity matrix in our case),  $\mathbf{w}_{k-1}$  is the process noise with covariance Q,  $z_k \in \mathbb{R}^6$  is the measurement vector, H is the observation matrix (also identity), and  $v_k$  is the measurement noise with covariance R.

**1279** The Kalman filter recursively estimates the state  $x_k$  by:

$$
\text{Prediction Step:} \quad \hat{\mathbf{x}}_{k|k-1} = \mathbf{F} \hat{\mathbf{x}}_{k-1|k-1},\tag{27}
$$

,

$$
\mathbf{P}_{k|k-1} = \mathbf{F} \mathbf{P}_{k-1|k-1} \mathbf{F}^{\top} + \mathbf{Q},\tag{28}
$$

Update Step: 
$$
\mathbf{K}_k = \mathbf{P}_{k|k-1} \mathbf{H}^\top (\mathbf{H} \mathbf{P}_{k|k-1} \mathbf{H}^\top + \mathbf{R})^{-1}
$$
, (29)

$$
\hat{\mathbf{x}}_{k|k} = \hat{\mathbf{x}}_{k|k-1} + \mathbf{K}_k(\mathbf{z}_k - \mathbf{H}\hat{\mathbf{x}}_{k|k-1}),
$$
\n(30)

$$
\mathbf{P}_{k|k} = (\mathbf{I} - \mathbf{K}_k \mathbf{H}) \mathbf{P}_{k|k-1},
$$
\n(31)

**1289 1290** where  $\hat{\mathbf{x}}_{k|k}$  is the estimated state at time k,  $\mathbf{P}_{k|k}$  is the estimated covariance, and  $\mathbf{K}_k$  is the Kalman gain.

**1291 1292** The initial state  $\hat{\mathbf{x}}_{0|0}$  is set to the first measurement, and the initial covariance  $\mathbf{P}_{0|0}$  is set to the identity matrix.

### **1293 1294** Multi-Level Downsampling

**1295** To reduce computational complexity, we create two additional levels of downsampled sequences using average pooling:

- Level-1: Downsampled by a factor of  $s_1$ .
- Level-2: Downsampled by a factor of  $s_1 \times s_2$ .

**1299 1300 1301** The downsampling is performed by averaging over non-overlapping windows of size  $s_i$ , for  $i = 1, 2$ . For example, for Level-1, the downsampled sequence  $S_1$  is obtained as:

$$
S_1[n] = \frac{1}{s_1} \sum_{k=(n-1)s_1+1}^{ns_1} S[k], \quad n = 1, 2, \dots, \left\lfloor \frac{L_S}{s_1} \right\rfloor, \tag{32}
$$

**1306** where  $L<sub>S</sub>$  is the length of the original sequence S.

### **1307 1308** Hierarchical Bias Search

**1309 1310** At each level, we perform a search for the optimal temporal bias  $b$  and matching length  $l$  that minimize the distance between the source and target sequences.

### **1311 1312** *Distance Metric*

**1313 1314** We define the distance between two sequences  $S$  and  $T$  over a matching window of length  $l$  as the mean Euclidean distance between their accelerometer and gyroscope data:

**1315 1316**

 $d_{\rm acc}(S, T; b, l) = \frac{1}{l}$  $\sum^l$  $k=1$  $\|\mathbf{a}_{S}[k + b] - \mathbf{a}_{T}[k]\|_{2}$  $(33)$ 

$$
d_{\rm gyr}(S,T;b,l) = \frac{1}{l} \sum_{k=1}^{l} ||\mathbf{g}_S[k+b] - \mathbf{g}_T[k]||_2, \qquad (34)
$$

**1322 1323 1324** where  $a_S[k]$  and  $g_S[k]$  are the accelerometer and gyroscope measurements of sequence S at time k, respectively.

#### **1325** *Coarse Search at Level-2*

**1326 1327** At the lowest resolution (Level-2), we perform a coarse search over a large range of biases b:

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 $b \in [b_{\min}, b_{\max}],$  (35)

**1331** where  $b_{\text{min}}$  and  $b_{\text{max}}$  are chosen based on the expected maximum temporal misalignment.

**1332 1333 1334** For each candidate bias b, we compute the distances  $d_{\text{acc}}$  and  $d_{\text{gyr}}$  and record the bias that minimizes these distances:

$$
b_{\rm acc}^{(2)} = \arg\min_b d_{\rm acc}(S_2, T_2; b, l_b),\tag{36}
$$

$$
b_{\text{gyr}}^{(2)} = \arg\min_{b} d_{\text{gyr}}(S_2, T_2; b, l_b),\tag{37}
$$

**1340 1341** where  $l_b$  is the matching length at bias b, determined by the overlapping length of the sequences after applying the bias.

#### **1342** *Refined Search at Level-1 and Level-0*

**1343 1344 1345 1346** Using the biases obtained at Level-2 as center points, we perform refined searches at higher resolutions (Level-1 and Level-0). The search ranges at each higher level are narrowed down around the biases found at the previous level:

$$
b_{\min}^{(i)} = b^{(i+1)} - \delta^{(i)},\tag{38}
$$

$$
b_{\max}^{(ii)} = b^{(i+1)} + \delta^{(i)}, \quad i = 1, 0,
$$
\n(39)

**1350 1351** where  $\delta^{(i)}$  is a small range that depends on the downsampling factor.

**1352** At each level, we update the biases:

$$
b_{\rm acc}^{(i)} = \arg\min_{b \in [b_{\rm min}^{(i)}, b_{\rm max}^{(i)}]} d_{\rm acc}(S_i, T_i; b, l_b),\tag{40}
$$

$$
b_{\rm gyr}^{(i)} = \arg\min_{b \in [b_{\rm min}^{(i)}, b_{\rm max}^{(i)}]} d_{\rm gyr}(S_i, T_i; b, l_b),\tag{41}
$$

**1359 1360** for  $i = 1, 0$ .

#### **1361** Optimal Bias and Alignment

**1363 1364** After performing the refined searches, we obtain the optimal biases  $b_{\text{acc}}^{(0)}$  and  $b_{\text{gyr}}^{(0)}$  at the original data level (Level-0). We choose the final bias  $b^*$  and matching length  $l^*$  based on the minimum distances:

$$
b^* = \text{median}(b_{\text{acc}}^{(0)}, b_{\text{gyr}}^{(0)}),
$$
  
\n
$$
l^* = \min(L_S - b^*, L_T),
$$
\n(42)

$$
\frac{1367}{1368}
$$

**1362**

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**1370** where  $L_S$  and  $L_T$  are the lengths of the source and target sequences, respectively.

**1371 1372 1373** The source and target sequences are then aligned by shifting the source sequence by  $b^*$  and taking the first  $l^*$  samples:

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**1399 1400**  $S_{\text{aligned}}[k] = S[k + b^*], \quad k = 1, 2, ..., l^*$  $(44)$ 

$$
T_{\text{aligned}}[k] = T[k], \quad k = 1, 2, \dots, l^*.
$$
 (45)

### **1378 1379** Algorithm Summary

- **1380** The overall algorithm can be summarized as follows:
	- 1. Apply Kalman filter to denoise the source and target IMU sequences.
		- 2. Downsample the sequences to create Level-1 and Level-2 versions.
		- 3. At Level-2, perform a coarse search over a wide range of biases to find initial estimates  $b_{\text{acc}}^{(2)}$ and  $b_{\text{gyr}}^{(2)}$ .
		- 4. At Level-1, perform a refined search around  $b^{(2)}$  to obtain  $b^{(1)}$ .
			- 5. At Level-0, perform a final refined search around  $b^{(1)}$  to obtain the optimal biases  $b_{\text{acc}}^{(0)}$  and  $b_{\rm gyr}^{(0)}.$
			- 6. Compute the final bias  $b^*$  and matching length  $l^*$ .
			- 7. Align the source and target sequences using  $b^*$  and  $l^*$ .

#### **1394** Implementation Details

**1395 1396 1397** In our implementation, we set the downsampling factors to  $s_1 = 10$  and  $s_2 = 10$ , resulting in Level-1 and Level-2 sequences downsampled by factors of 10 and 100, respectively.

**1398** The search ranges at each level are defined as:



- **1402** Level-1:  $b \in [b^{(2)} - 10s_1, b^{(2)} + 10s_1],$  (47)
- **1403** Level-0:  $b \in [b^{(1)} - 10s_0, b^{(1)} + 10s_0],$  (48)

**1404 1405** where  $s_0 = 1$  is the downsampling factor at Level-0 (original data).

#### **1406** Computational Efficiency

**1407 1408 1409 1410 1411** By employing the multi-level hierarchical search, we significantly reduce the computational complexity compared to an exhaustive search at the original sampling rate. At Level-2, the coarse search over a wide range of biases is feasible due to the reduced sequence length. The refined searches at higher resolutions are limited to small ranges around the biases found at lower levels, ensuring that the total computational cost remains manageable.

# **1412** Visualization of the Alignment Results

**1413 1414 1415 1416** Fig. [11,](#page-27-0) Fig. [12](#page-28-0) and Fig. [11](#page-27-0) showcase the IMU registration results for two trajectories. The high degree of overlap between the two IMU streams after alignment demonstrates the effectiveness of our proposed method.

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### **1418** F MORE VISUALIZATION RESULTS

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#### **1420** More Examples on Our SEE-600K Dataset

**1421 1422 1423 1424 1425 1426** The additional visualizations provided in Fig. [14](#page-30-0) and Fig. [15](#page-31-0) demonstrate the diversity of the SEE-600K dataset. The dataset captures a wide variety of scenes, both indoors and outdoors, including objects like plants, buildings, and everyday items. This diversity reflects common realworld scenarios, ensuring comprehensive coverage of typical environments. The images span different lighting conditions, showcasing the dataset's ability to handle various illumination levels, from low to high light.

# **1427 1428** More Visualization on SEE-600K Dataset

**1429 1430 1431 1432 1433** Fig[.16,](#page-32-0)[17](#page-33-0)[,18](#page-34-0)[,19](#page-35-0)[,20,](#page-36-0)[21](#page-37-0) showcase additional visual results on the SEE-600K dataset. These examples further demonstrate the robustness and consistency of our proposed SEE-Net method. Notably, when using a brightness prompt of 0.5, SEE-Net is capable of generating more stable and higher-quality images. In some cases, the output even surpasses the quality of the ground truth normal-light image (GT), showing the strength of our approach in various lighting conditions.

**1434 1435 1436 1437 1438 1439** Additionally, it's important to highlight certain challenging cases, as shown in Fig. [20.](#page-36-0) For instance, in regions with highly detailed textures or areas requiring high-resolution recovery, all current methods, including ours, struggle to achieve optimal results. Despite this, SEE-Net continues to show relatively better performance compared to existing methods, particularly in maintaining image quality and stability. These results illustrate the potential of our method to handle complex scenarios, but they also indicate areas where further improvements could be made in future research.

**1440 1441 1442** By highlighting both the strengths and limitations of our approach, these visualizations provide valuable insights into the practical capabilities of SEE-Net across a wide range of real-world lighting conditions and complex scenes.

#### **1443** More Visualization on SDE Dataset

**1444 1445 1446 1447 1448 1449** Fig. [23](#page-39-0) and [24](#page-40-0) present additional visualizations from the SDE dataset, specifically focusing on challenging low-light outdoor scenes. These low-light environments often come with significant noise, which poses a substantial challenge for current low-light enhancement methods. Our method demonstrates stable performance in addressing these noisy scenes, effectively enhancing the image quality while mitigating the noise, thereby highlighting the robustness of our approach in handling complex low-light conditions.

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