# FRACTAL CALIBRATION FOR LONG-TAILED OBJECT DETECTION

Anonymous authors

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Paper under double-blind review

## Abstract

Real-world datasets follow an imbalanced distribution, which poses significant challenges in rare-category object detection. Recent studies tackle this problem by developing re-weighting and re-sampling methods, that utilise the class frequencies of the dataset. However, these techniques focus solely on the frequency statistics and ignore the distribution of the classes in image space, missing important information. In contrast to them, we propose Fractal CALibration (FRA-CAL): a novel post-calibration method for long-tailed object detection. FRACAL devises a logit adjustment method that utilises the fractal dimension to estimate how uniformly classes are distributed in image space. During inference, it uses the fractal dimension to inversely downweight the probabilities of uniformly spaced class predictions achieving balance in two axes: between frequent and rare categories, and between uniformly spaced and sparsely spaced classes. FRACAL is a post-processing method and it does not require any training, also it can be combined with many off-the-shelf models such as one-stage sigmoid detectors and two-stage instance segmentation models. FRACAL boosts the rare class performance by up to 8.6% and surpasses all previous methods on LVIS dataset, while showing good generalisation to other datasets such as COCO, V3Det and Open-Images. We provide the code in the Appendix.

## 1 INTRODUCTION

031 In recent years, there have been astonishing developments in the field of object detection Carion 032 et al. (2020); Chen et al. (2022); Lyu et al. (2022). Most of these works utilise vast, balanced, 033 curated datasets such as ImageNet1k Deng et al. (2009), or MS-COCO Lin et al. (2014) to learn 034 efficient image representations. However, in the real world, data are rarely balanced, in fact, they follow a long-tailed distribution Liu et al. (2019). When models are trained with long-tailed data, they perform well for the frequent classes of the distribution but they perform inadequately for the 037 rare classes Wang et al. (2020); Ren et al. (2020); Li et al. (2020). This problem poses significant 038 challenges to the safe deployment of detection and instance segmentation models in real-world safecritical applications such as autonomous vehicles, medical applications, and industrial applications, scenarios where rare class detection is paramount. 040

041 Many approaches address the long-tailed detection problem by employing adaptive re-weighting or 042 data resampling techniques to handle imbalanced distributions Wang et al. (2021a;b); Zang et al. 043 (2021). However all these methods require training. In contrast, in long-tailed image classification, 044 alternative methods focus on mitigating class imbalance during inference through a post-calibrated softmax adjustment (PCSA) Alexandridis et al. (2023); Ren et al. (2020); Hong et al. (2021). PCSA boasts strong performance, good compatibility with many methods like data augmentation, masked 046 image modeling, contrastive learning, and does not necessitate specialized loss function optimiza-047 tion, making it more user friendly Xu et al. (2023); Cui et al. (2021); Zhu et al. (2022). However, 048 current PCSA methods utilise solely the train set's class frequency  $p_s(y)$  as shown in Figure 1-left and they overlook the significance of the classes' dependence on the location distribution  $p_s(y, u)$ . This is a significant limitation of previous PCSA methods because the location information is a 051 critical indicator considering the correlation between classes y and their respective locations u. 052

Motivated by the class-location dependence Kayhan & Gemert (2020), in this work, we investigate a novel way to incorporate location information into post-calibration for imbalanced object detection



Figure 1: Previous PCSA used the class prior  $p_s(y)$  to align the learned source distribution  $p_s(y, u|x)$ with the balanced target distribution  $p_t(y|x)$ , without considering the space information u, highlighted in blue. FRACAL embeds space information  $p_s(y, u)$  into class calibration, via the fractal dimension and aligns the learned  $p_s(y, u|x)$  with  $p_t(y, u|x)$  better than previous works.

to boost the performance of rare classes by fully exploiting dataset statistics. We empirically show that naively injecting location statistics results in inferior performance because the location information is sparse for the rare classes. To overcome this, we propose FRACAL (FRActal CALibration), a novel post-calibration method based on the fractal dimension, as shown in Figure 1-right. Our method aggregates the location distribution of all objects in the training set, using the box-counting method Schroeder (2009). This resolves the sparsity problem and significantly enhances the performance of both frequent and rare classes as shown in our experiments.

Our method comes with several advantages. First, it performs an effective class calibration, suitable for the object detection task, using the dataset's class frequencies. Secondly, it captures the class-location dependency Kayhan & Gemert (2020), using the fractal dimension, and it fuses this information into class calibration. This results in a better and unique space-aware logit-adjustment technique that complements the frequency-dependent class calibration method and achieves higher overall performance compared to previous PCSA techniques.

**FRACAL** can be easily combined with both one-stage and two stage detectors, Softmax and Sigmoid-based models, various instance segmentation architectures, various backbones, sampling strategies, and largely increase the performance during the inference step. FRACAL significantly advances the performance on the challenging LVISv1 benchmark Gupta et al. (2019) with no training, or additional inference cost by 8.6% rare mask average precision  $(AP_r^m)$ .

- 084 Our **contributions** are as follows:
  - For the first time, we show the importance of the class-location dependence in postcalibration for long-tailed object detection.
  - We capture the location-class dependence via a space-aware long-tailed object detection calibration method based on the fractal dimension.
  - Our method performs remarkably on various detectors and backbones, on both heavily imbalanced datasets such as LVIS and less imbalanced datasets such as COCO, V3DET and OpenImages, outperforming the state-of-the-art by up to 8.6%.
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# 2 RELATED WORK

096 General Object Detection. General object detection Redmon & Farhadi (2017); Ren et al. (2015); 097 Lin et al. (2017b); Liu et al. (2016); Carion et al. (2020); Zhu et al. (2021); Sun et al. (2021); Chen 098 et al. (2022); Li et al. (2022e) and instance segmentation He et al. (2017); Huang et al. (2019); Cai 099 & Vasconcelos (2019); Chen et al. (2019a); Wang et al. (2019); Bolya et al. (2019); Li et al. (2022e) have witnessed tremendous advancements. Recently, transformer-based detectors were proposed 100 which use self-attention to directly learn object proposals Carion et al. (2020); Zhu et al. (2021), 101 or diffusion-based methods which use a de-noising process to learn bounding boxes Chen et al. 102 (2022) and segmentation masks Gu et al. (2022b). However, all of these methods struggle to learn 103 the rare classes when trained with long-tailed data Gupta et al. (2019); Oksuz et al. (2020) due to 104 the insufficient rare samples. To this end, FRACAL enhances the rare class performance using a 105 space-aware logit adjustment that can be easily applied during inference. 106

**Long-tailed image classification.** In the past years, the long-tailed image recognition problem has received great attention, as demonstrated by many recent surveys Oksuz et al. (2020); Zhang et al.

FRACAL (ours)

 $\begin{array}{c} \text{Adjustment} \\ z'_y = z_y - \tau \log(p_s(y)) \\ z'_y = -z_y \cdot \log(p_s(y)) \\ z'_y = z_y - \log(p_s(y)) + \log(p_t(y)) \\ p'_y = \frac{p_y/n_y^{\gamma}}{p_{bg} + \sum p_y/n_y^{\gamma}}, y \notin bg \\ z'_y = \frac{z_y - (\mu_y - \min_y(\mu_y))}{\varsigma_y}, y \notin bg \\ z'_y = \mathbf{S}(\mathbf{C}(z_y)) / \sum_{j=1}^{C+1} \mathbf{S}(\mathbf{C}(z_y)) \end{array}$ Method Dependency Log. Adj. Menon et al. (2021) Frequency IIF Alexandridis et al. (2023) Frequency PC-Softmax Hong et al. (2021) Frequency Norcal Pan et al. (2021) Frequency LogN Zhao et al. (2022a) Frequency

Space & Frequency

to past works, FRACAL considers both frequency and space statistics as shown in Section 3.

Table 1: Post-calibration techniques in long-tailed tasks.  $\tau$  and  $\gamma$  are hyper-parameters, bg is the

background class,  $\mu_{y}$  and  $\varsigma_{y}$  are estimated class mean and standard deviation respectively. Compared

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121 (2023b); Yang et al. (2022a) and newly created benchmarks Yang et al. (2022b); Tang et al. (2022); 122 Gu et al. (2022a). In long-tailed classification, the works could be split into two groups, represen-123 tation learning and classifier learning. Representation learning techniques aim to efficiently learn 124 rare class features using oversampling Park et al. (2022); Hong et al. (2022); Zang et al. (2021), 125 contrastive learning Li et al. (2022d); Zhu et al. (2022); Cui et al. (2023), using ensemble or fusion models Wang et al. (2021c); Li et al. (2022c;b); Cui et al. (2022); Aimar et al. (2023), knowledge 126 distillation Li et al. (2022c); He et al. (2021); Li et al. (2021a), knowledge transfer Liu et al. (2019); 127 Parisot et al. (2022); Zhu & Yang (2020), sharpness aware minimisation Zhou et al. (2023a;b); Ma 128 et al. (2023) and neural collapse Li et al. (2023); Zhong et al. (2023); Liu et al. (2023). Classifier 129 learning techniques aim to adjust the classifier in favour of the rare classes via decoupled train-130 ing Kang et al. (2020); Zhang et al. (2021b); Hsu et al. (2023), margin adjustment Menon et al. 131 (2021); Ren et al. (2020); Hong et al. (2021); Cao et al. (2019); Hyun Cho & Krähenbühl (2022); 132 Zhao et al. (2022b); Alexandridis et al. (2023); Ye et al. (2020) and cost-sensitive learning Cui et al. 133 (2019); Khan et al. (2017); Wang et al. (2017). Among these works, the Post-Calibrated Softmax 134 Adjustment (PCSA) method Menon et al. (2021); Hong et al. (2021); Ma et al. (2024) distinguishes 135 itself through both its strong performance and the absence of any training requirements. However, 136 most of the classifier and representation learning techniques are hard to adopt in long-tailed object 137 detection. This difficulty arises from the larger imbalance inherent in this task, amplified by the presence of the background class Mullapudi et al. (2021); Yang et al. (2022a). Moreover, the opti-138 misation of models for this task becomes more complex due to multiple sources of imbalance such as 139 batch imbalance, class imbalance and task imbalance as outlined in this survey Oksuz et al. (2020). 140 For this reason, we develop FRACAL, which is a post-calibration method tailored to the long-tailed 141 object detection task. Different from post-calibration classification methods Menon et al. (2021); 142 Hong et al. (2021), FRACAL enhances the detection performance by leveraging class-dependent 143 space information derived from the fractal dimension. Through space-aware logit-adjustment, FRA-144 CAL mitigates biases in both the detection's location and classification axes. 145

Long-tailed object detection. The most prevalent technique is adaptive rare class re-weighting, 146 which could be applied using either the statistics of the mini-batch Hsieh et al. (2021); Tan et al. 147 (2020); Wang et al. (2021b) or the statistics of the gradient Tan et al. (2021); Li et al. (2022a). Other 148 works use adaptive classification margins based on the classifier's weight norms Wang et al. (2022); 149 Li (2022), classification score Feng et al. (2021); He et al. (2022); Wang et al. (2021a), activation 150 functions Alexandridis et al. (2022; 2024), group hierarchies Li et al. (2020); Wu et al. (2020) and 151 ranking loss Zhang et al. (2023a). Many works use data resampling techniques Zang et al. (2021); 152 Gupta et al. (2019); Kang et al. (2020); Feng et al. (2021); Wu et al. (2020) or external rare class augmentation Zhang et al. (2022; 2021a). All these works optimise the model on the long-tailed 153 distribution and require the construction of a complicated and cumbersome training pipeline. In 154 contrast, our method operates during the model's inference stage thus it is easier to use and less 155 evasive to the user's codebase. 156

157 Norcal Pan et al. (2021) was the first method to apply a post-calibration technique in imbalanced ob-158 ject detection, achieving promising results without training the detector. They proposed to calibrate 159 only the foreground logits using the train-set's label statistics and applied a re-normalisation step. LogN Zhao et al. (2022a) proposed to use the model's own predictions to estimate the class statistics 160 and applied standardisation in the classification layer. However LogN, requires forward-passing the 161 whole training set through the model to estimate the weights, thus it is slower than FRACAL, which



172 Figure 2: During imbalanced object detection, the model makes more frequent class predictions 173 like *hat* and less rare class predictions like *tiara* both of which have strong upper location bias. FRACAL utilises fractal dimension and debiases the logits both in the frequency and space axes, 174 making fewer hat predictions and more *tiara* predictions that are evenly spread, achieving space and 175 frequency balance and increasing performance. 176

is not model-dependent. Also, both methods do not utilise the spatial statistics of the classes which 177 are valuable indicators since the classes and their location are correlated Kayhan & Gemert (2020). 178 To this end, FRACAL balances the detectors using both class and space information, largely sur-179 passing the performance of the previous methods. FRACAL can be easily combined with two-stage 180 softmax-based models like MaskRCNN He et al. (2017), or one-stage sigmoid detectors such as 181 GFLv2 Li et al. (2021b) achieving great results without training or additional inference cost. 182

**Relation to previous works.** In Table 1, we contrast our work to previous post-calibration methods 183 used in classification and object detection. As the Table suggests, all prior methods are frequencydependent and none of them considers the space statistics. 185

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#### 3 METHODOLOGY

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189 In Subsection 3.1, we pose the problem of calibration for classification; in Subsection 3.2, we extend 190 it to the problem of object detection and we analyse the location-class dependence. We then, in Subsection 3.3, capture class-dependent space information via the fractal dimension and in Subsection 3.4, we combine it with the class-calibration method and extend it for binary object detectors. We 192 show our approach in Fig. 2.

#### 3.1 BACKGROUND: CLASSIFICATION CALIBRATION 195

196 Let  $f_u(x;\theta) = z$  be a classifier parameterised by  $\theta$ , x the input image, y the class, z the logit, 197  $\bar{y}$  is the model's prediction and  $p_s(y)$  and  $p_t(y)$  the class priors on the train and test distributions respectively. The post-calibration equation is: 199

$$\bar{y} = \arg\max_{w} (f_y(x;\theta) + \log(p_t(y)) - \log(p_s(y))).$$
(1)

201 This has been numerously analysed in previous literature Menon et al. (2021); Alexandridis et al. 202 (2023); Hong et al. (2021); Ren et al. (2020); Lipton et al. (2018) and we derive it in Appendix. In short, this shows that to get better performance, one can align the model's predictions with the test 203 distribution, by subtracting  $\log(p_s(y))$  and adding  $\log(p_t(y))$  in the logit space. We now extend it 204 to object detection. 205

#### 3.2 CLASSIFICATION CALIBRATION FOR OBJECT DETECTION 207

208 In classification, p(y) can be easily defined using the dataset's statistics, by using instance frequency 209  $n_y$ , i.e.  $p(y) = \frac{n_y}{\sum_{i=1}^{C} n_i}$ . In object detection, this is not the case because p(y) is affected by the 210 location and the object class. Accordingly, we define the class priors as: 211

$$p(y, o, u) = p(y|o, u) \cdot p(o, u) = p(y, u) \cdot p(o, u),$$
(2)

213 where o denotes the generic object occurrence and u is the location inside the image. By substituting 214 Eq. 2 inside Eq. 1, we get  $\bar{y}$  as:

$$\bar{y} = \arg\max_{y} (f_y(x;\theta) + \log(p_t(y,u) \cdot p_t(o,u)) - \log(p_s(y,u) \cdot p_s(o,u))).$$
(3)



Figure 3: Different grid sizes affect the object distribution estimation. When the grid is coarse, e.g.,  $1 \times 1$  or  $2 \times 2$ , there is no or little location information. When it is finer, e.g.,  $128 \times 128$ , the probability is sparse, giving noisy estimates for the rare classes.

The term p(o, u) in Eq. 3 cannot be calculated apriori as it depends on the model's training (e.g., the 229 IoU sampling algorithm, how the object class is encoded etc<sup>1</sup>). Despite this,  $p_s(o, u) \approx p_t(o, u)$ , 230 as we show in the Appendix, which means that the object distributions of the train and the test set remain the same and only the foreground class distribution changes. As a result:

$$\bar{y} = \arg\max_{y} (f_y(x;\theta) + \log(p_t(y,u)) - \log(p_s(y,u)))$$
(4)

Next, we show how the location parameter u affects Eq. 4.

#### 3.2.1 LOCATION-CLASS INDEPENDENCE.

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First, we consider the scenario where the location u does not give any information. In this scenario, u and y are independent variables, thus  $p(y, u) = p(y) \cdot p(u)$  and we rewrite Eq. 4 as:

$$\bar{y} = \arg\max_{y} (f_y(x;\theta) + \log(p_t(y) \cdot p_t(u)) - \log(p_s(y) \cdot p_s(u)))$$
  
= 
$$\arg\max_{y} (f_y(x;\theta) + \log(p_t(y)) - \log(p_s(y))),$$
(5)

where p(u) is the probability of a random location in the image space and it has been simplified

because it is the same in both source and target distributions, i.e.,  $p_s(u) = p_t(u)$ .

247 In theory  $p_t(y)$  is unknown, thus Eq.5 cannot be applied. Despite that, we found that setting  $p_t(y) =$ 248  $\frac{1}{C}$  works well, because it forces the model to do balanced detections on the test set. In practice, this 249 maximises average precision because this metric independently evaluates all classes and it rewards 250 balanced detectors Everingham et al. (2010). Accordingly, the Classification (C) calibration of the logit  $z_y$  is: 251

$$\mathbf{C}(z_y) = \begin{cases} z_y - \log_\beta(\frac{n_y}{\sum_i^C n_i}) + \log_\beta(\frac{1}{C}), & y \in \{1, ..., C\} \\ z_y, & y = \mathbf{bg}, \end{cases}$$
(6)

where  $\beta$  is the base of the logarithm that we optimise through hyperparameter search. The back-255 ground logit remains unaffected because of the assumption that the object distribution is the same 256 in train and test set  $p_s(o, u) \approx p_t(o, u)$ , (this assumption is also found in previous works Pan et al. 257 (2021); Zhao et al. (2022a)). 258

To this end, Eq. 6 can get good performance as shown in our ablation study but it is limited because 259 the assumption that  $p(y, u) = p(y) \cdot p(u)$  is not correct. In the real world, the object detection distri-260 bution has a strong center bias, as shown in Fig.3 and discussed in Oksuz et al. (2020). Furthermore, 261 the location is correlated with the class Kayhan & Gemert (2020), therefore,  $p(y, u) \neq p(y) \cdot p(u)$ . 262 As we show, the location provides valuable information for the long-tailed detection task and we 263 enhance Eq. 6 by fusing location information. 264

#### 265 3.2.2 LOCATION-CLASS DEPENDENCE. 266

267 One way to compute p(y, u) is by counting the class occurrences  $n_u(\mathbf{u})$  along locations that fall 268 inside the cell  $\mathbf{u} = [i, j]$  as shown in Fig. 3-left. To do so, we discretise the space of various image

<sup>&</sup>lt;sup>1</sup>Typically object detectors use an extra background logit bg to implicitely learn p(o, u).



Figure 4: a) An example of the box counting method for the class *cow*. It iteratively counts the boxes containing its center  $\nu$ , as G grows. b-c) The blue points are all  $G - \nu$  pairs, out of them only the orange points are used to calculate the slope  $\Phi$  based on the quadratic rule  $G = \lfloor \sqrt{n_y} \rfloor$ . d-e) Scatter-plot of  $\Phi$  against instance frequency, there is a weak correlation i.e. 0.35 for LVISv1 and 0.375 for COCO using Pearson's correlation.

resolutions into a normalised square grid  $U_{G \times G}$  of fixed size  $G \in \mathbb{N}$  and count class occurrences inside every grid cell. Accordingly, the grid dependent calibration is defined as:

$$\mathbf{C}_{G}(z_{y,\mathbf{u}}) = \begin{cases} z_{y,\mathbf{u}} - \log_{\beta}(p_{s}(y,\mathbf{u})) + \log_{\beta}(p_{t}(y,\mathbf{u})) \\ z_{y,\mathbf{u}}, & if \quad y = \mathbf{bg}, \end{cases}$$
(7)

where  $z_{y,\mathbf{u}}$  is the predicted proposal whose center falls inside the discrete cell  $\mathbf{u} = [i, j]$  and  $p_t(y, \mathbf{u})$ is uniform, i.e.,  $p_t(y, \mathbf{u}) = \frac{1}{C} \cdot \frac{1}{G^2}$ .

However, the choice of the grid size G largely affects the estimation of p(y, u), as shown in Fig. 300 3-right. For example, if we use smaller G, the generic object distribution becomes denser and little location information is encoded. If we use larger G, the distribution becomes sparse. This is problematic for the rare classes because they are already sparse and their location information is noisy. In Table 4-e, we show that this method shows limited performance.

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#### 3.3 CALIBRATION USING FRACTALS

To solve the sparsity problem introduced by the grid-size, we use fractal dimension  $\Phi$  Panigrahy et al. (2019), which is a metric independent of the grid size G. To calculate  $\Phi$ , we use the boxcounting method Schroeder (2009):

$$\Phi(y) = \lim_{G \to \infty} \frac{\log \sum_{j=0}^{G-1} \sum_{i=0}^{G-1} \mathbb{1}(n_y(\mathbf{u}))}{\log(G)},$$
(8)

where 1 is the indicator function. For objects in 2D images, as in our case,  $\Phi(y) \in [0, 2]$ , where 0 is only one object, 1 shows that the objects lie across a line and 2 shows that they are located uniformly across the image space.

For brevity, we rewrite  $\nu_y = \sum_{j=0}^{G-1} \sum_{i=0}^{G-1} \mathbb{1}(n_y(\mathbf{u}))$  and we give an example in Fig. 4-a. In practice, Eq. 8 cannot be computed because by increasing *G*, the computation becomes intractable. Instead, we approximate  $\Phi$ , by evaluating nominator-denominator pairs of Eq. 8 for various values of *G* up to a threshold *t* and then fit a line to those points. The slope of this line approximates  $\Phi(y)$ , because it considers all computed  $G - \nu_y$  pairs.

**Dealing with rare classes.** To select the threshold t, we use the quadratic rule  $G \le t = \lfloor \sqrt{n_y} \rfloor$ . The motivation for this rule is simple, for example, if an object is rare, e.g., it appears 4 times in the whole training set, then it can, at most, fill a grid of size  $2 \times 2$ . For objects with fewer occurrences we cannot compute  $\Phi$  and thus we assign  $\Phi = 1$ . Using this rule, we define the maximum number of pairs that are required for fitting the "fractality" line highlighted in orange in Fig. 4-b and Fig. 4-c. For example, the rare object *birdbath* appears 12 times in the training set, thus we use the first three orange points in Fig. 4-c that correspond to  $G = \{1, 2, 3\}$ , to fit the "fractality" line, resulting in a large  $\Phi = 1.67$ . This rule ensures that the fractal dimension computation does not underestimate the rare classes and it gives robust measurements that increase rare class performance as shown in our experiments. For the *cow* object that has larger frequency we use more  $G - \nu$  orange pairs to fit the line as shown in Fig. 4-b, resulting in  $\Phi = 1.80$ .

331 **Relationship to frequency.** As shown in Fig. 4-d, the fractal dimension weakly correlates with 332 frequency for the LVISv1 dataset, i.e., 0.35 using Pearson correlation. However, there are many rare 333 classes with large  $\Phi \approx 2$ , showing that our threshold selection technique is robust for small sample 334 sets. Also, the correlation for the COCO dataset in Fig. 4-e is similar to LVIS because these datasets have the same images. This shows that the relationship of the estimated fractal dimension and the 335 frequency is dependent on the image data itself and not the class imbalance and it highlights that our 336 method is robust for different label distributions, it is not a purely frequency-based method, and it 337 captures the class space statistics effectively. 338

Usage. After calculating  $\Phi$  for all classes in the training set, we perform Space-dependent class calibration (S) during inference:

$$\mathbf{S}(z_y) = \begin{cases} \frac{\sigma(z_y)}{\Phi(y)^{\lambda}}, & y \in \{1, ..., C\}\\ \sigma(z_y), & y = \mathrm{bg}, \end{cases}$$
(9)

where  $\sigma(z_y) \in (0, 1)$  is the model's prediction for class y, with  $\sigma()$  the Softmax activation, and  $\lambda \ge 0$  is a hyperparameter. Eq. 9 downweighs the classes that appear most uniformly and it upweighs the classes that appear less uniformly. In practice, this scheme enforces a centre bias for frequent classes and no bias for rare classes, as shown in Fig. 2-bottom-right. Intuitively, removing the bias from the rare classes is better than rectifying it because it produces balanced detectors and aligns better with the target distribution as shown in our ablation and our qualitative results.

350 3.4 LOCALISED CALIBRATION

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By putting Eq. 6 and Eq. 9 together, we get the final FRActal CALibration (FRACAL) as:

$$\operatorname{FRACAL}(z_y) = \frac{\operatorname{S}(\operatorname{C}(z_y))}{\sum_{j=1}^{C+1} \operatorname{S}(\operatorname{C}(z_j))}.$$
(10)

Our proposed method tackles the classification imbalance using additional space statistics. On the classification axis, we use the class priors  $p_s(y)$  and perform logit adjustments. On the space axis, we use the fractal dimension  $\Phi(y)$  to perform a space-aware calibration that accounts for the object's location distribution  $p_s(y, u)$ . In Eq. 10, we renormalise both foreground and background logits to preserve a probabilistic prediction after the space calibration in Eq. 9.

**Extending to binary classifiers.** In long-tailed object detection there are many works that use only binary classifiers Alexandridis et al. (2022); Tan et al. (2020; 2021); Li et al. (2022a); Wang et al. (2021b); Hyun Cho & Krähenbühl (2022); Hsieh et al. (2021). In this case, the logit  $z_i$  performs two tasks simultaneously: It discriminates among the foreground classes and performs background-toforeground classification. Thus, to correctly apply foreground calibration, we first need to decouple the foreground and background predictions. To do so, we filter out the background proposals using the model's predictions as follows:

$$\operatorname{FRACAL}_{b}(z_{i}) = \eta(\operatorname{C}(z_{i}) - \log_{\beta}(\frac{\Phi(y)^{\lambda}}{\sum_{i}^{C} \Phi(i)^{\lambda}}) + \log_{\beta}(\frac{1}{C})) \cdot \eta(z_{i}), \tag{11}$$

where  $\eta(z_i)$  is the sigmoid activation function that acts as a filter for low-scoring proposals. Compared to Eq. 10, Eq. 11 performs class calibration and space calibration in logit space, lowering the false-positive detection rate.

4 Results

#### 375 4.1 EXPERIMENTAL SETUP

We use the Large Vocabulary Instance Segmentation (LVISv1) dataset Gupta et al. (2019) which consists of 100k images in the train set and 20k images in the validation set. This dataset has 1,203

81	Method	Arch.	$AP^{m}$	$AP_r^m$	$AP_c^m$	$AP_{f}^{m}$	$AP^b$
82	NorCal Pan et al. (2021)		25.2	19.3	24.2	29.0	26.1
83	LogN Zhao et al. (2022a)		27.5	21.8	27.1	30.4	28.1
84	GOL Alexandridis et al. (2022)	<b>D</b> 50	<u>27.7</u>	21.4	27.7	30.4	27.5
85	ECM Hyun Cho & Krähenbühl (2022)	KJU	27.4	19.7	27.0	<u>31.1</u>	27.9
86	CRAT w/ LOCE Wang et al. (2024)		27.5	21.2	26.8	31.0	<u>28.2</u>
87	FRACAL (ours)		28.6	23.0	$2\bar{8}.\bar{0}$	31.5	28.4
88	NorCal Pan et al. (2021)		27.3	20.8	26.5	31.0	28.1
89	LogN Zhao et al. (2022a)		<u>29.0</u>	<u>22.9</u>	28.8	31.8	29.8
90	GOL Alexandridis et al. (2022)		<u>29.0</u>	22.8	29.0	31.7	29.2
91	ECM Hyun Cho & Krähenbühl (2022)	R101	28.7	21.9	27.9	<u>32.3</u>	29.4
02	ROG Zhang et al. (2023a)		28.8	21.1	<u>29.1</u>	31.8	28.8
000	CRAT w/ LOCE Wang et al. (2024)		28.8	22.0	28.6	32.0	29.7
593	FRACAL (ours)		29.8	24.5	29.3	32.7	29.8
202	(******)	1				- /-	

Table 2: Comparison against SOTA on LVISv1 dataset. All competing methods use MaskRCNN
 and our method reaches the best results in all metrics. † denotes our re-implementation with RFS.

classes grouped according to their image frequency into *frequent* (those that contain > 100 images), 396 *common* (those that contain  $10 \sim 100$  images) and *rare* classes (those that contain < 10 images) in 397 the training set. For evaluation, we use average mask precision  $AP_m$ , average box precision  $AP_b$ and  $AP_f^m$ ,  $AP_c^m$  and  $AP_r^m$  that correspond to  $AP^m$  for frequent, common and rare classes. Unless 398 mentioned, we use Mask R-CNN He et al. (2017) with FPN Lin et al. (2017a), ResNet50 He et al. 399 (2016), repeat factor sampler (RFS) Gupta et al. (2019), Normalised Mask and cosine classifier as 400 used in Wang et al. (2021a), CARAFE Wang et al. (2019) and we train the baseline model using the 401 2x schedule He et al. (2019), SGD, learning rate 0.02 and weight decay 1e-4. For Swin-T, we train 402 the baseline model with the 1x schedule, RFS, AdamW Kingma & Ba (2014) and 0.001 learning 403 rate. During inference, we set the IoU threshold at 0.3 and the mask threshold at 0.4.FRACAL is 404 applied before the non-maximum suppression step. We use the mmdetection framework Chen et al. 405 (2019b) and train the models using V100 GPUs.

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## 4.2 MAIN RESULTS

**Comparison to SOTA.** In Table 2, we compare FRACAL to the state-of-the-art using ResNet50 and ResNet101. Using ResNet50, FRACAL significantly surpasses GOL Alexandridis et al. (2022) by 0.9pp  $AP^m$  and by 1.6pp  $AP_r^m$ . On ResNet101 FRACAL achieves 29.8%  $AP^m$  and 24.5%  $AP_r^m$ , outbesting GOL by 0.8pp and 1.7pp respectively.

FRACAL achieves excellent results not only for rare categories but also for frequent ones, due to the use of fractal dimension, which allows the model to upscale the predictions of frequent but non-uniformly located categories. It achieves  $31.5\% AP_f^m$  with ResNet50 and  $32.7\% AP_f^m$  with ResNet101 and surpasses the next best method, ECM Hyun Cho & Krähenbühl (2022) by 0.4pp.

417 418 Compared to the previous post-calibration method, Norcal Pan et al. (2021), FRACAL increases per-419 formance by 3.4pp  $AP^m$ , 3.7pp  $AP^m_r$ , 3.8pp  $AP^m_c$ , 2.5pp  $AP^m_f$  and 2.3pp  $AP^b$  using ResNet50. 420 This is because FRACAL boosts both rare and frequent categories via classification and space calibration, respectively, while Norcal only boosts the rare categories and lacks space information.

We also compare our method in Transformer backbones. Using Swin-T, FRACAL considerably outperforms Seesaw Wang et al. (2021a) by 1.2pp  $AP^m$ , 1.7pp  $AP^m_r$ , 1.2pp  $AP^m_c$ , 1.0pp  $AP^m_f$ and 0.8pp  $AP^b$  as shown in Table 3-a. Using Swin-S, FRACAL largely surpasses Seesaw in all metrics and particularly in  $AP^m_r$  by 2.2pp which is a significant 8.6% relative improvement for the rare classes.

427 Results on object detectors. We evaluate FRACAL with common object detectors in Table 3428 b using ResNet50. FRACAL boosts the overall and rare category performance of both one-stage
429 detectors such as ATSS Zhang et al. (2020) or GFLv2 Li et al. (2021b) and two-stage detectors such
430 as Cascade RCNN Cai & Vasconcelos (2019) and APA-MaskRCNN Alexandridis et al. (2024).
431 Note that on sigmoid-detectors such as ATSS or GFLv2, FRACAL largely boosts the performance of rare and common categories but it slightly reduces the performance of frequent categories. Since

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Table 3: In (a), we show that FRACAL can be used with Swin transformers Liu et al. (2021) and
surpass the SOTA. In (b), we show that FRACAL can be used with both Sigmoid and Softmax based
detectors and improve their precision.

						Method	$AP^{b}$	$AP_r^b$	$AP_c^b$	$AP_f^b$
Method	$ AP^{m} $	$AP_r^m$	$AP_c^m$	$AP_{f}^{m}$	$AP^{b}$	ATSS Zhang et al. (2020)	25.3	15.8	23.4	31.6
RFS-(T)	27.7	17.9	27.9	31.8	27.1	w/ FRACAL (ours)	26.7	20.8	25.9	30.9
Seesaw-(T)	29.5	24.0	29.3	32.2	<u>29.5</u>	GFLv2 Li et al. (2021b)	26.6	14.7	25.1	33.5
GOL-(T)	28.5	21.1	<u>29.5</u>	30.6	28.3	w/ FRACAL (ours)	28.2	19.4	27.2	33.2
FRACAL-(T)	30.7	25.7	30.5	33.2	30.3	GFLv2 (DCN) Li et al. (2021b)	27.4	13.7	26.1	34.8
RFS-(S)	30.9	21.7	31.0	34.7	31.0	w/ FRACAL (ours)	28.9	18.7	27.9	34.5
Seesaw-(S)	<u>32.4</u>	<u>25.6</u>	<u>32.8</u>	<u>34.9</u>	<u>32.9</u>	APA Alexandridis et al. (2024)	26.9	14.3	26.2	33.2
GOL-(S)	31.5	24.1	32.3	33.8	32.0	w/ FRACAL (ours)	29.2	22.1	28.0	33.7
FRACAL-(S)	33.6	27.8	33.9	35.9	33.4	C-RCNN Cai & Vasconcelos (2019)	28.6	16.5	27.8	34.9
						w/ FRACAL (ours)	31.5	24.3	31.0	35.3
(a) Results using Swin (T/S) and MaskRCNN. (b) Comparisons using various detectors and ResNet5									let50.	

Table 4: Ablations using MaskRCNN-ResNet50. C and S denote the class and location calibration.

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	С	S	$AP^m$	$AP_r^m$	λ	AF	m	$AP_r^m$	$AP_c^m$	$AP_f^m$	$AP^{b}$		0	ranc	lom	]	RFS	
_			22.8	8.2	0.0	28	.0	22.4	27.3	31.2	27.4		β	$AP^m$	$AP_r^m$	$AP^{m}$	$AP_r^n$	n
		$\checkmark$	25.6	13.7	1.0	28	.5	23.0	28.0	31.6	28.3		2	19.9	14.7	19.9	18.8	
	$\checkmark$		26.3	16.5	2.0	28	.0 5	23.0	28.0	31.5	28.4		e	25.1	16.6	25.8	21.1	
	$\checkmark$	$\checkmark$	27.3	19.0	4.0	28	.5	23.2 23.4	27.9	31.3	28.4		10	26.3	16.5	28.0	22.4	
(:	(a) Random sampler.			pler.	(b) <i>i</i>	Åbla	atio	n study	y of $\beta$	, with	RFS.		(c) A	blation	stud	y of $\lambda$ .		
	С	S	$AP^{m}$	$AP_r^m$	Meth	bod	$\Delta P^m$	$AP^{m}$	$\Delta P^m$	$AP^m$	$AP^b$							
_			25.7	15.8		1	20.0	22.4	27.2	21.2	27.4			Method	1	$AP^m$	$AP_n^m$	$AP^{b}$
		$\checkmark$	27.7	20.7	G= G=	2	28.0	17.5	27.5	31.2	27.4		Inv	ert FRAG	CAL	27.4	20.5	26.9
	$\checkmark$		28.0	22.4	G=	4	25.0	10.5	25.4	31.1	24.9		Nor	mal FRA	CAL	28.6	23.0	28.4
	$\checkmark$	$\checkmark$	28.6	23.0	oui	rs	28.6	23.0	28.0	31.5	28.4							
(	(d) Results using RFS.				(e) I	Resi	ilts	with C	Grid-b	ased n	nethod	ł.	(f) In	vert FF	RACA	L is in	ferior.	

the sigmoid activation performs independent classification, the binary version of FRACAL struggles to properly calibrate the predicted unnormalised vector. This limitation was also found in previous works Pan et al. (2021) which also reported that binary logit adjustment produces performance tradeoffs between frequent and rare categories. For softmax-based detectors, such as Cascade RCNN and APA, FRACAL boosts all categories.

## 462 4.3 ABLATION STUDY AND ANALYSIS

The effect of each module. FRACAL consists of simple modules that we ablate in Table 4-a. First, MaskRCNN with CARAFE Wang et al. (2019), normalised mask predictor Wang et al. (2021a), cosine classifier Wang et al. (2021a) and random sampler achieves 22.8%  $AP^m$  and 8.2% rare category  $AP_r^m$ . On top of this, the fractal dimension calibration (S) improves  $AP^m$  and  $AP_r^m$  by 2.8pp and 5.5pp respectively.

Using only the classification calibration, (C),  $AP^m$  and  $AP_r^m$  are enhanced by 3.5pp and 8.3pp respectively, because this technique majorly upweights the rare classes. When (S) is added, then it further increases  $AP^m$  by 1.0pp and  $AP_r^m$  by 2.5pp compared to only (C), reaching 27.3%  $AP^m$ and 19.0%  $AP_r^m$ . This suggests that (S) is useful and the detector can benefit from space information. The same trend is observed with RFS in Table 4-d, however, both calibration methods have lower gains because RFS partly balances the classes via oversampling.

**Class calibration parameter search.** We further ablate the choice of the log base  $\beta$  in Eq. 6, using the most common cases: 2 (bit), *e* (nat), and 10 (hartley). As shown in Table 4-b, the base-10 is the best as it achieves 26.3%  $AP^m$  and 16.5%  $AP_r^m$  with the random sampler and 28.0%  $AP^m$  and 22.4%  $AP_r^m$  with RFS, thus we use it for all experiments on LVIS. We also observe that further increasing  $\beta$  does not come with a performance improvement.

**Fractal dimension coefficient.** We ablate the choice of the  $\lambda$  coefficient in the fractal dimension calibration Eq. 9. As shown in Table 4-c, the optimal performance is achieved with  $\lambda = 2$  which increases the rare categories by 0.6pp, the common categories by 0.7pp, the frequent categories by 0.3pp, the overall mask performance by 0.6pp and the box performance by 1.0pp.

484 **Comparison to grid-dependent calibration.** We compare FRACAL against the grid-based 485 method, Eq. 7, in Table 4-c. When G = 1 the method does not consider any location information because all predictions fall inside the same grid cell. This achieves the best performance and it

Method	$AP^m$	$AP^b$	
ResNet50 He et al. (2016	)	35.4	39.4
with FRACAL (ours)		35.8	39.9
SE-ResNet50 Hu et al. (20)	18)	36.9	40.5
with FRACAL (ours)		37.4	41.1
CB-ResNet50 Woo et al. (20	)18)	37.3	40.9
with FRACAL (ours)		37.8	41.5
Swin-T Liu et al. (2021)		41.6	46.0
with FRACAL (ours)	41.9	46.4	
(a) Results on COCO with M	laskR	CNN.	
Method	$AP^{b}$	$AP_{50}^b$	$AP_{75}^b$
APA Alexandridis et al. (2024)	29.9	37.6	32.9
with FRACAL (ours)	30.3	37.7	33.2
(b) Results on V3Det Wang	et al.	(2023) w	ith
FasterRCNN and Re	esNet	50.	
Method		Detector	$AP_{50}^b$
CAS Liu et al. (2020)	EDCNN	65.0	
CAS with FRACAL (ours)	FREIM	67.0	
CAS Liu et al. (2020)			66.3
ECM Hyun Cho & Krähenbühl (2	022)	CRCNN	65.8
CAS with FRACAL (ours)			67.5

(c) Results on OpenImages using ResNet50.





Figure 5: Detection results in LVIS-val. FRACAL detects more uniformly located rare classes in (2c) and less uniformly located frequent ones in (3c) than the baseline in (2b) and (3b).

507 is the same result with the  $\lambda = 0$  of Table 4-c. When the grid size G is enlarged, the performance of 508 the rare classes drops significantly because the estimated prior distribution  $p_s(y, \mathbf{u})$  becomes sparse (see Fig. 3). FRACAL does not suffer from this problem, because it re-weights all proposals based 509 on fractal dimension. 510

511 **Inverting FRACAL.** We further examine the opposite weighting logic, which is to upweight the 512 uniform located classes and downweight the non-uniform located classes. This technique further 513 rectifies the location bias instead of removing it from the object detectors. As Table 4-f shows, the 514 Invert FRACAL method is inferior to the normal one, because it produces an imbalanced detector.

515 Generalisation to other datasets. We test FRACAL on MS-COCO Lin et al. (2014), V3DET Wang 516 et al. (2023) and OpenImages Kuznetsova et al. (2020) to understand its generalisation ability. The 517 first two datasets are fairly balanced therefore, we do not expect our long-tailed designed detector to 518 massively outperform the others. In Table 5 a-b, FRACAL increases the performance of all models, by an average of 0.5pp  $AP^b$  and  $AP^m$  on COCO and by 0.4pp  $AP^b$  on V3DET. In Table 5-c, we 519 520 show that FRACAL outperforms ECM using CascadeRCNN by 1.7pp and it further increases the performance of CAS by 2.0pp and 1.2pp using FasterRCNN and CascadeRCNN respectively. 521

522 **Qualitative Analysis.** In Fig. 5, we show: (a) the ground truth distribution, (b) the baseline and 523 (c) FRACAL predicted distributions concerning general objects (1), the rare class *ferret* (2) and the 524 frequent class zebra (3). FRACAL achieves better precision than the baseline because it predicts 525 fewer generic objects in (1-c) than the baseline (1-b); it predicts more rare classes that are more uniformly located in (2-c) than the baseline in (2-b); and it predicts less frequent classes that have a 526 stronger center-bias as shown in (3-c) than the baseline in (3-b). These results show that FRACAL 527 aligns its predictions better with the ground-truth distribution than the baseline. 528

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#### 5 CONCLUSION

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We propose FRACAL, a novel post-calibration method for long-tailed object detection. Our method 534 performs a space-aware logit adjustment, that utilises the fractal dimension and incorporates space information during calibration. FRACAL majorly boosts the performance of the detectors and makes 536 more rare class predictions that are evenly spread inside the image. We show that FRACAL can be easily combined with both one-stage Sigmoid detectors and two-stage Softmax segmentation models. Our method boosts the performance of detectors by up to 8.6% without training or additional 538 inference cost, surpassing the SOTA in the LVIS benchmark and showing good generalisation to COCO, V3Det and OpenImages.

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## A BACKGROUND: CLASSIFICATION CALIBRATION

We theoretically derive the classification calibration for image classification. Let  $p_s(y|x)$  and  $p_t(y|x)$  be the source and target conditional distributions. Using the Bayes theorem, we write the source and target conditional distributions as:

$$p_s(y|x) = \frac{p_s(x|y)p_s(y)}{p_s(x)}, p_t(y|x) = \frac{p_t(x|y)p_t(y)}{p_t(x)}$$
(12)

Big Dividing them, we write the target conditional distribution:

$$p_t(y|x) = \frac{1}{\kappa(x)} \frac{p_t(y)}{p_s(y)} p_s(y|x) \frac{p_t(x|y)}{p_s(x|y)}$$
(13)

where  $\kappa(x) = \frac{p_t(x)}{p_s(x)}$ . During training, we approximate  $p_s(y|x)$  by model  $f_y(x;\theta) = z$  and a scorer function  $s(x) = e^x$  for multiple category classification. Thus, the learned source conditional distribution is  $p_s(y|x) \propto e^{f_y(x;\theta)}$ . Substituting it inside Eq. 13, we rewrite the target condition distribution as:

$$p_t(y|x) \propto \frac{1}{\kappa(x)} \frac{p_t(y)}{p_s(y)} e^{f_y(x;\theta)} \frac{p_t(x|y)}{p_s(x|y)}$$

$$= d(x,y) \cdot e^{f_y(x;\theta) + \log(p_t(y)) - \log(p_s(y)) - \log(\kappa(x))}$$
(14)

where we assume that  $d(x, y) = \frac{p_t(x|y)}{p_s(x|y)} = 1$ . This is a reasonable assumption, in cases where both train and test generating functions come from the same dataset, as it is in our benchmarks. In inference, we calculate the prediction  $\bar{y}$  by taking the maximum value of Eq. 14:

$$\bar{y} = \arg\max_{y} e^{(f_y(x;\theta) + \log(p_t(y)) - \log(p_s(y)) - \log(\kappa(x)))}$$

$$= \arg\max_{y} (f_y(x;\theta) + \log(p_t(y)) - \log(p_s(y)))$$
(15)

where  $\kappa(x)$  is simplified because it is a function of x and it is invariant to  $\arg \max_y$ . Eq. 15 is the post-calibration method Menon et al. (2021); Hong et al. (2021). It can be used during inference to achieve balanced performance by injecting prior knowledge inside the model's predictions, via  $p_t(y)$  and  $p_s(y)$ , in order to align the source with the target label distribution and compensate for the label shift problem.

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## **B** FRACTAL DIMENSION VARIANTS

We explore various ways for computing the fractal dimension using the box-counting method Schroeder (2009), the information dimension Rényi (1959) (Info), and a smooth variant (Smooth-Info). The information variant is defined as:

Info-
$$\Phi(y) = \lim_{G \to \infty} \frac{\log \sum_{j=0}^{G-1} \sum_{i=0}^{G-1} \frac{\mathbb{1}(n_y(\mathbf{u}))}{G}}{\log(G)}$$
 (16)

864	Dimension	$AP^{m}$	$AP_r^m$	$AP^{b}$
865	Info	28.6	23.2	28.3
866	SmoothInfo	28.6	23.4	28.3
867	Box	28.6	23.0	28.4
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Table 6: Fractal Dimension Variants using MaskRCNN with ResNet50 and RFS on LVISv1. All of the are robust and we have chosen the Box variant in the main paper.

871 It is the similar to the box-counting dimension, except for the box count which is normalised by 872 dividing by the grid size G. This way, the information dimension is represented by the growth rate 873 of the probability  $p = \frac{1(n_y(\mathbf{u}))}{G}$  as G grows to infinity.

In practise, the quantity  $\mathbb{1}(n_y(\mathbf{u}))$  can be frequently zero for many locations  $\mathbf{u}$  especially for rare classes that have few samples and are sparsely located. For this reason, we also proposed a smooth information variant defined as:

Smooth-
$$\Phi(y) = \lim_{G \to \infty} \frac{1 + \log \sum_{j=0}^{G-1} \sum_{i=0}^{G-1} \frac{1 + 1(n_y(\mathbf{u}))}{G}}{\log(G)}$$
 (17)

This Equation is inspired by the smooth Inverse Document Frequency Robertson (2004) used in natural language processing and its purpose is to smooth out zero values in  $1(n_y(\mathbf{u}))$  calculation.

All variants are robust and SmoothInfo achieves slightly better  $AP_r^m$  because its calculation is more tolerant to few samples compared to the box-counting method. However, SmoothInfo and Info achieve slightly worse  $AP^b$ , thus we use the box-counting method in the main paper.

# C OBJECT DISTRIBUTIONS

We show that the object distribution  $p_s(o, u)$  in the training set is similar to the object distribution  $p_t(o, u)$  on the test set in the LVIS v1 dataset Gupta et al. (2019). As shown in Figure 6, the distributions are close therefore we can safely assume that  $p_s(o, u) \approx p_t(o, u)$ . This explains the reason why the background logit should remain intact during calibration because there does not exist label shift for the generic object class (also for the background class) between the train and test sets.



Figure 6: Comparison between the  $p_s(o, u)$  (left) and  $p_t(o, u)$  (right) in LVISv1 dataset. The distributions are similar, therefore we can safely assume that  $p_s(o, u) \approx p_t(o, u)$ .